

# ATLAS muon NSW

Conception, construction, completion and current status

A quick history from the beginning

27.10.2023

T. Kawamoto

# Upgrade of small wheels → NSW



A major 'phase-1' upgrade

Installed during LS2

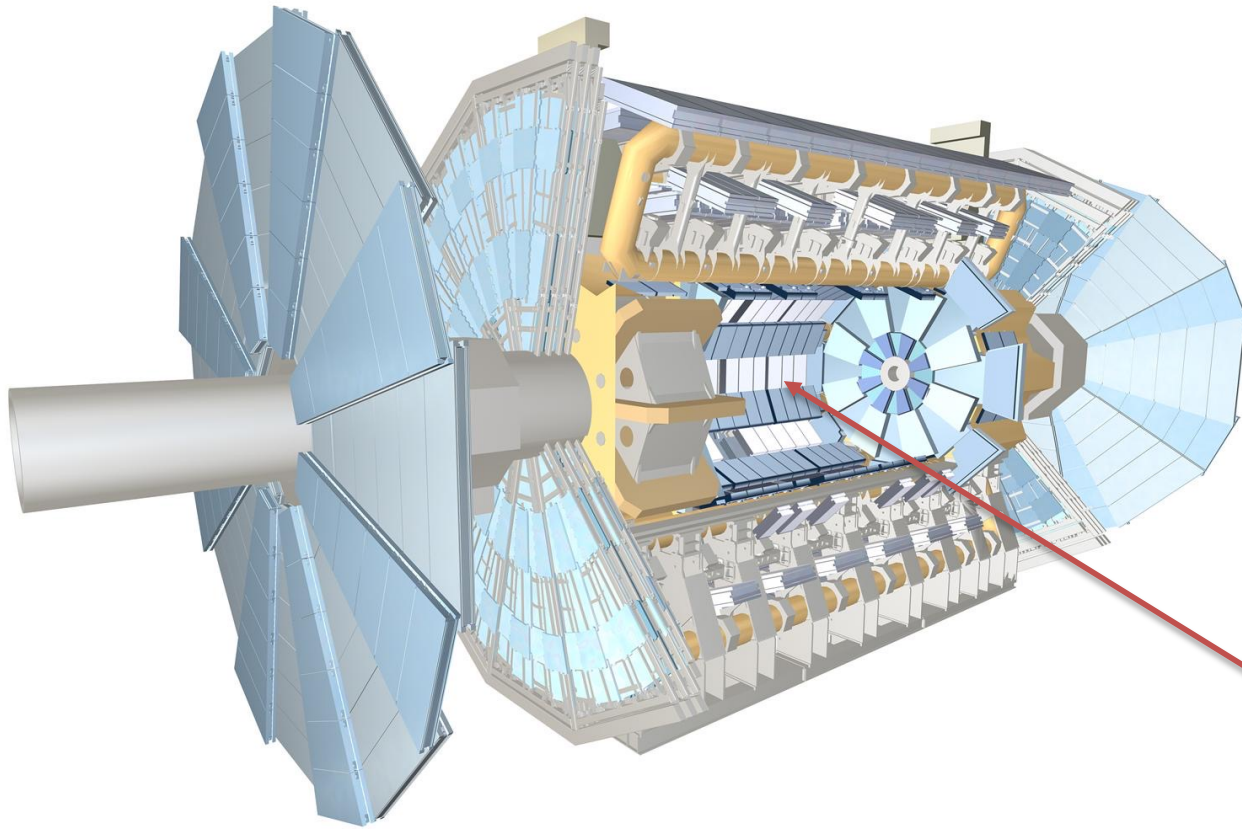
Started operation from 2022

## A long time ago in a galaxy ....

- Discussions/studies of ATLAS upgrade have been around since  $\sim 2004$ , well before the completion of ATLAS detector and start of pp collisions.
- This is in the context of LHC luminosity upgrade, which is inspired by the fact that the lifetime (radiation damage) of the triplet (focusing) magnets comes after several  $100 \text{ fb}^{-1}$ .
- New triplets will be designed for replacement, and at the same time allowing much higher instantaneous luminosity:  $\sim 5 \times 10^{34}$ . sLHC  $\rightarrow$  HL-LHC
- Main interest of ATLAS (and CMS) was inner tracker.
- Muon system did not have specific idea of upgrade. Some worry about increase of **cavern background** : background hit rate  $\rightarrow$  performance degradation (MDT).
  - Some R&Ds started looking for muon detector for high rate
- Serious Muon studies started after the collision data have been collected (essentially from 2010): hit rate measurements, and some surprise.



# Cavern background



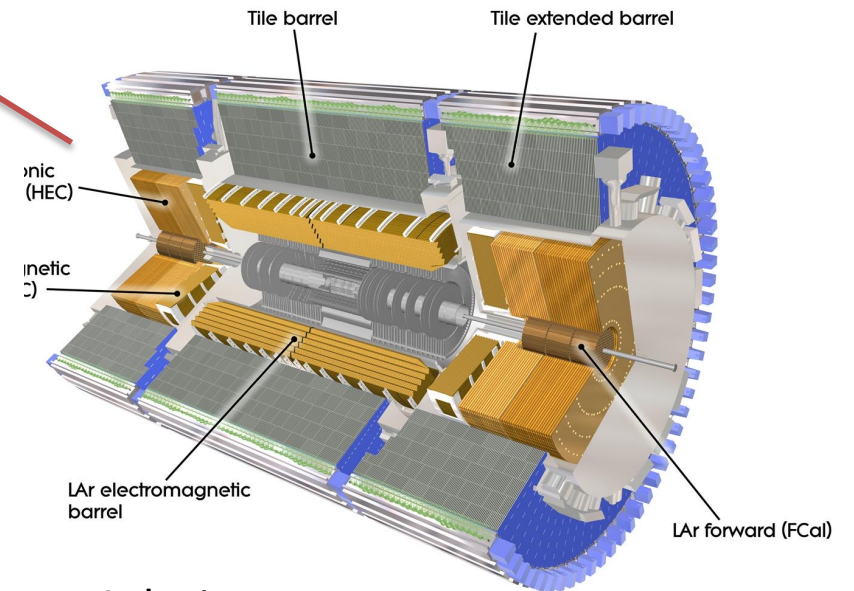
Muon system

The space outside of the calorimeter is filled by huge number of particles (mostly low energy).

- Leak from calorimeters
- Leak from beam pipe shieldings

The muon system is irradiated by the “cavern background” particles

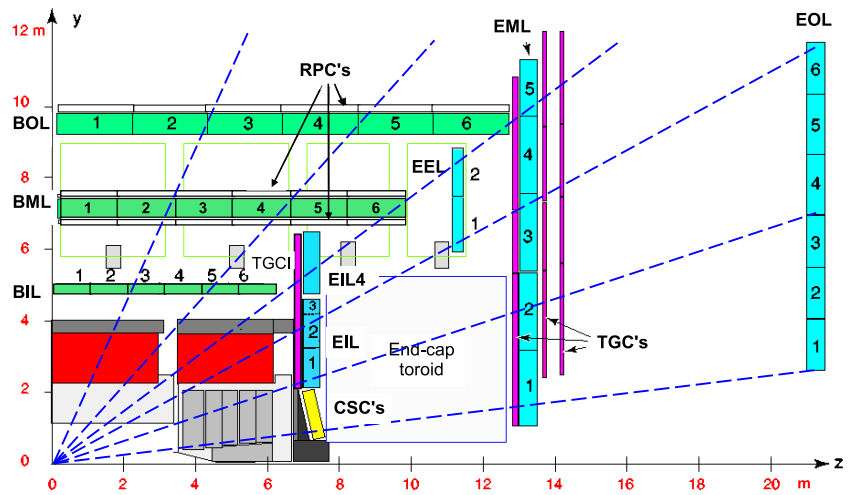
- Background hits
- Radiation damage



Calorimeters



# ATLAS muon detector original

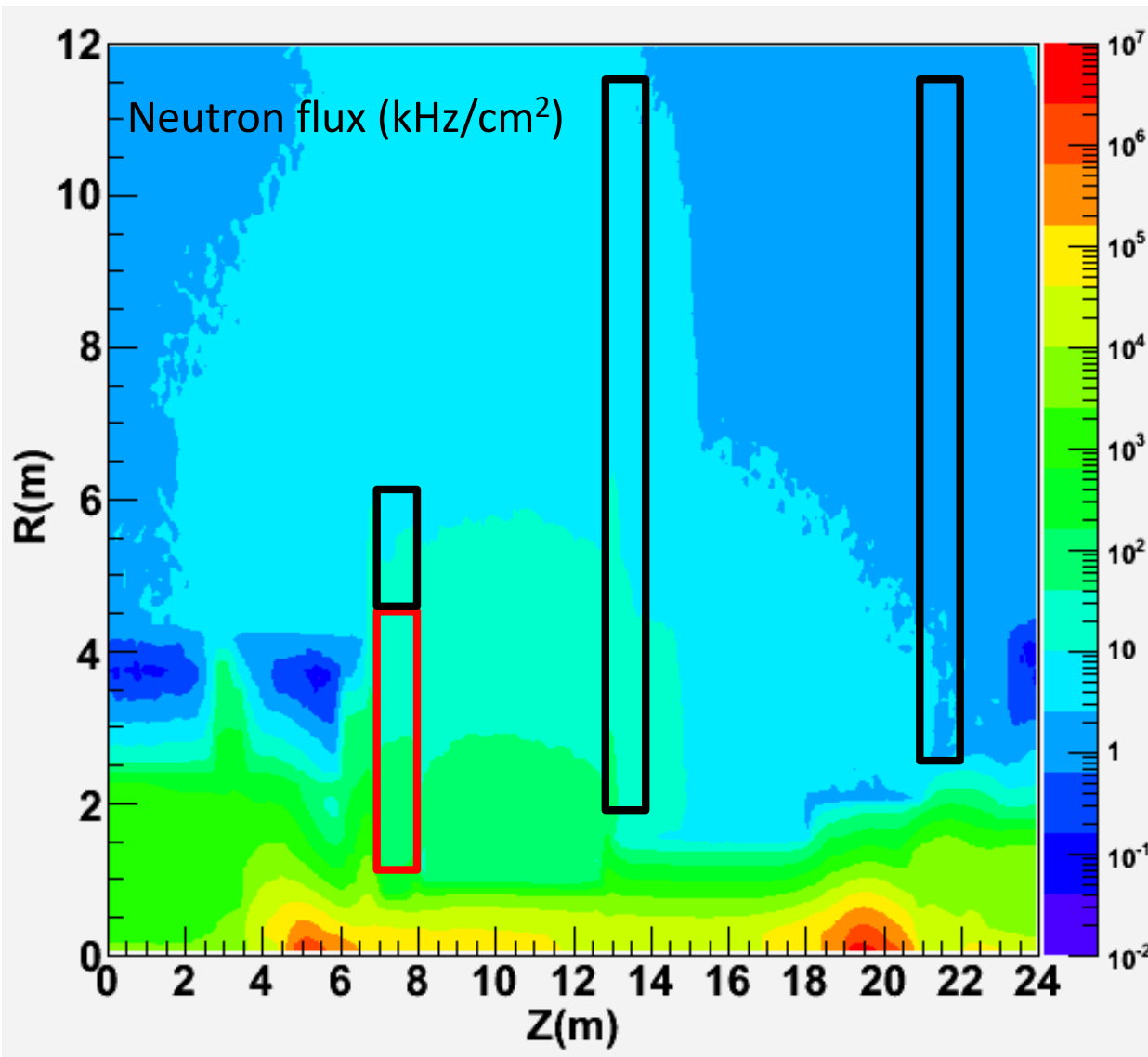


27.10.2023



Tatsuo Kawamoto

# Cavern background



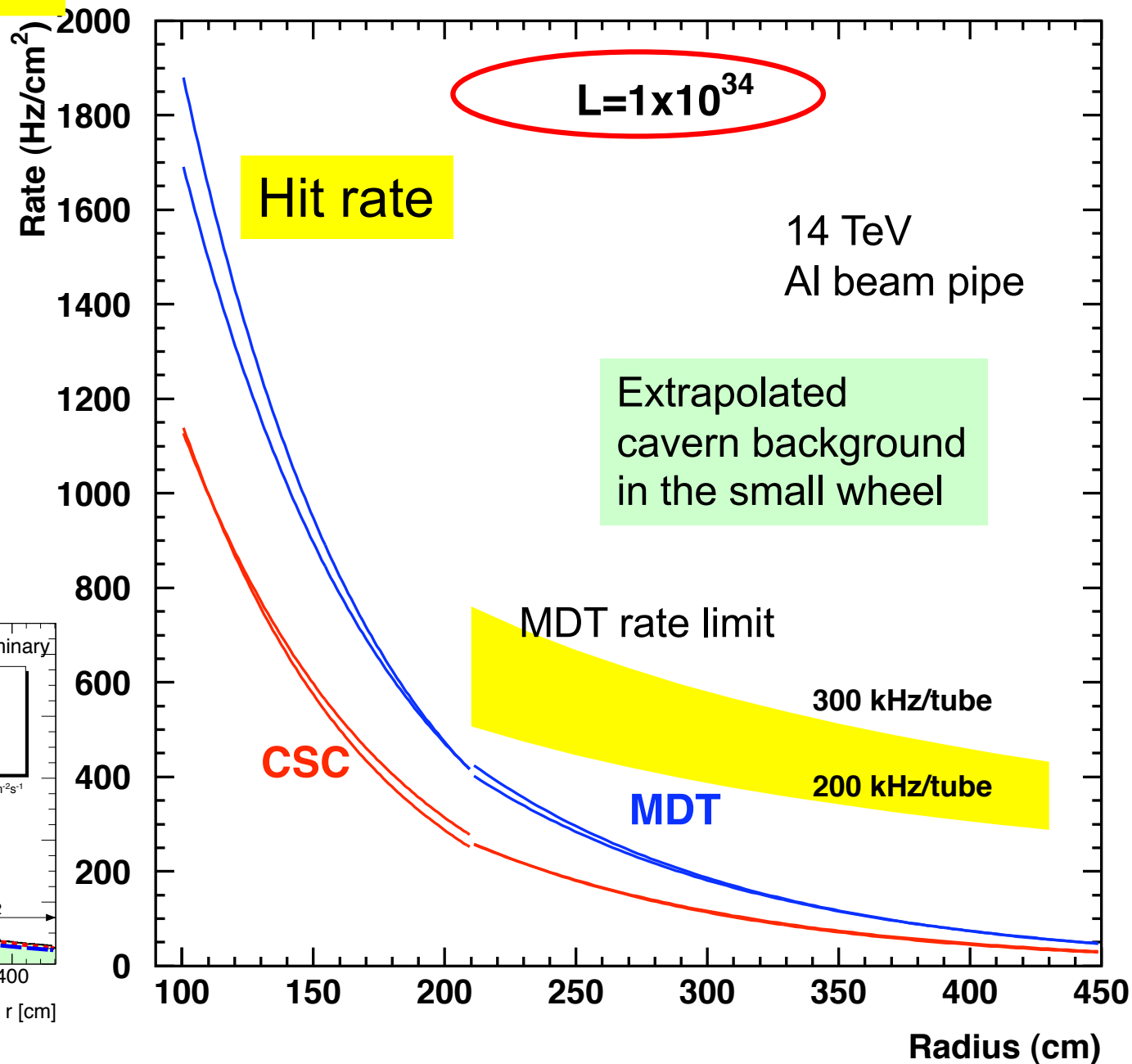
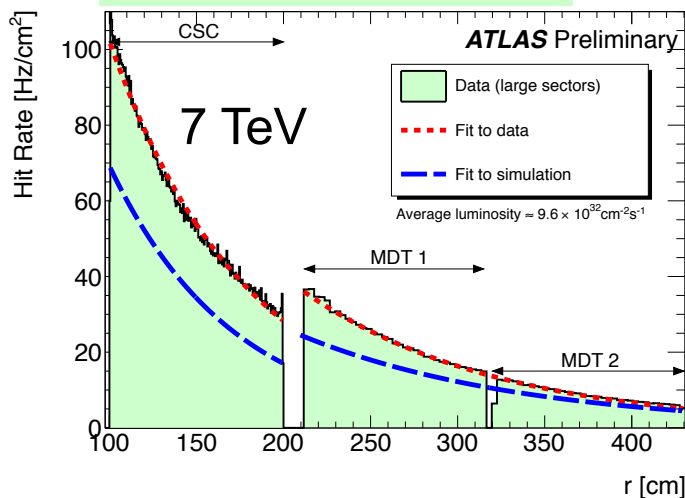
Mostly low energy photons, and neutrons  
+ much lower level of charged particle flux.

Muon detector and electronics were designed based on the estimation (simulation available at that time) + safety factor (typically x5), but for the design luminosity ( $1 \times 10^{34}$ ).

# Cavern background

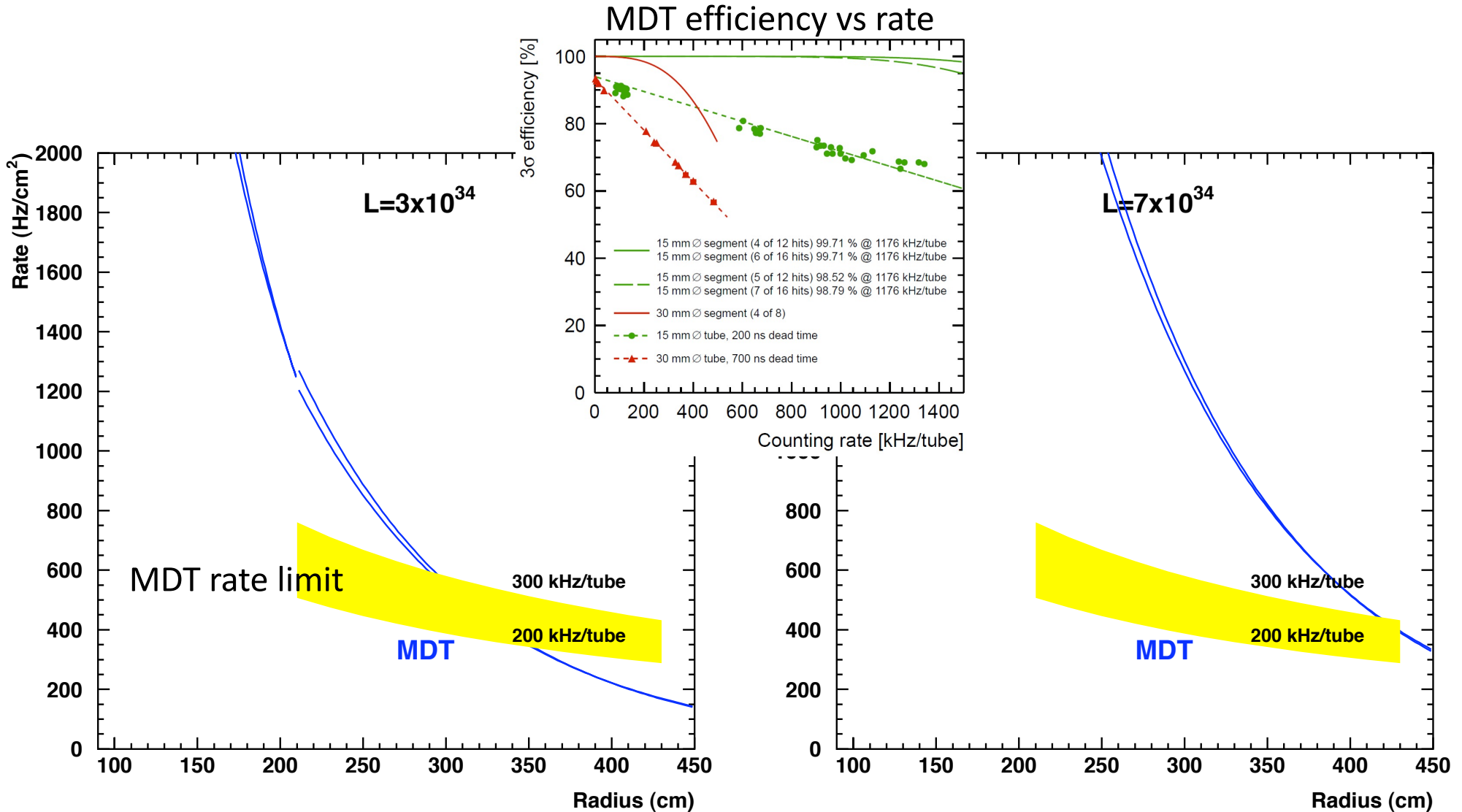
$1 \times 10^{34}$  is OK, but little margin for higher lumi

Measured cavern background in the small wheel





# MDT cavern background issue



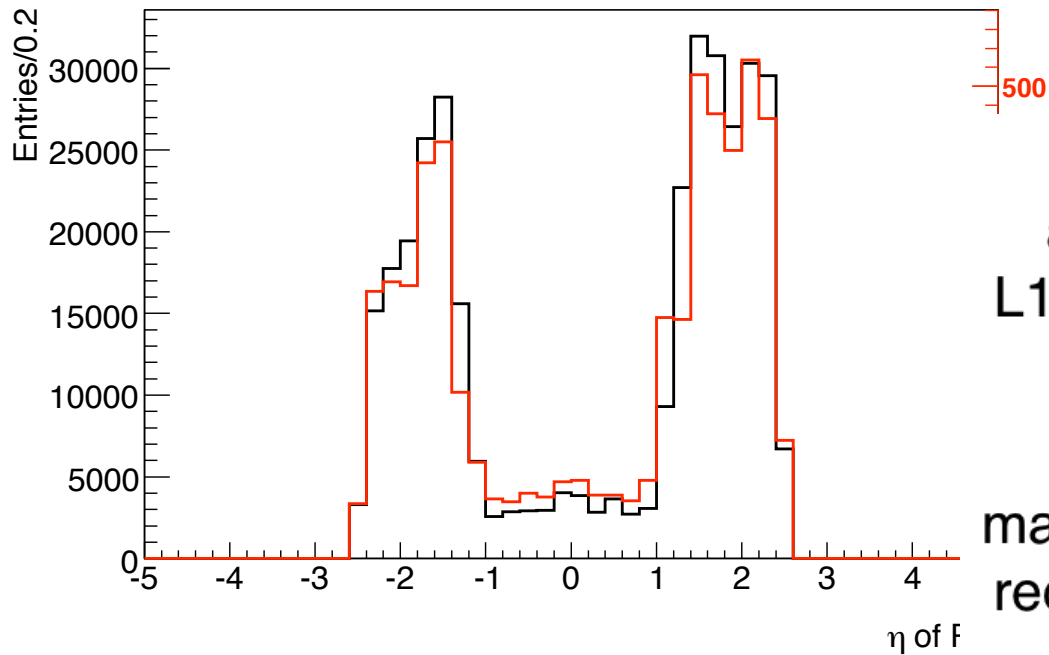
MDT (and perhaps CSC) of Small Wheel will need upgrade for high luminosity.

# More serious problem : Muon L1 issue

Background trigger will saturate the L1 bandwidth at high lumi

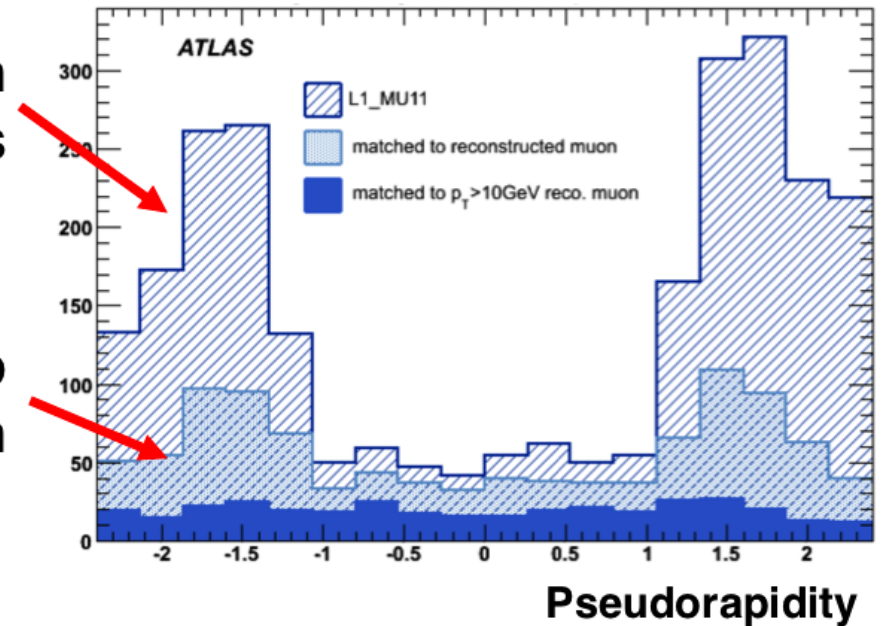
MU20 vs  $\eta$  2010/2011 data

50 ns interval



all muon  
L1 triggers

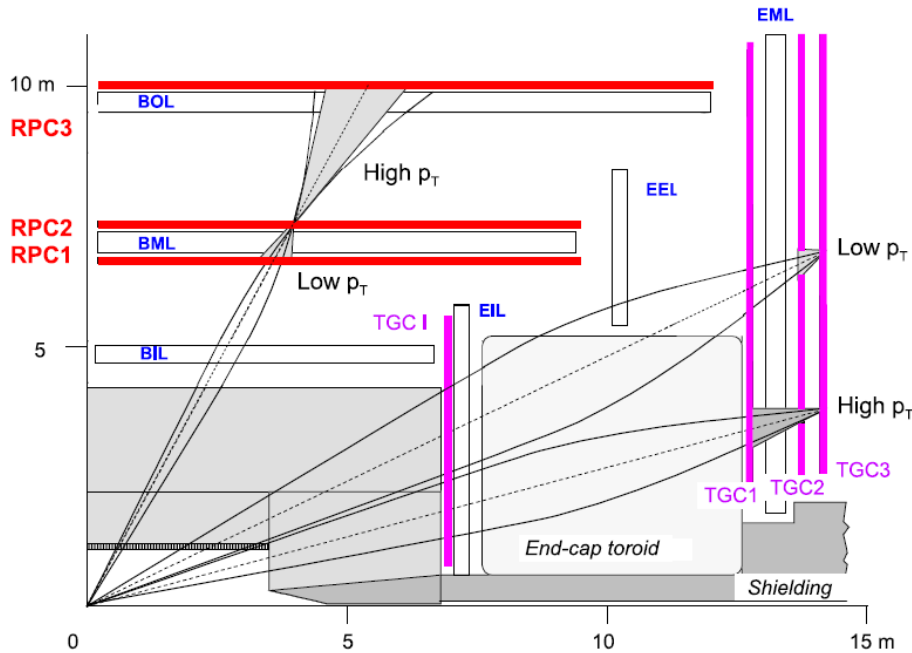
matched to  
reco muon



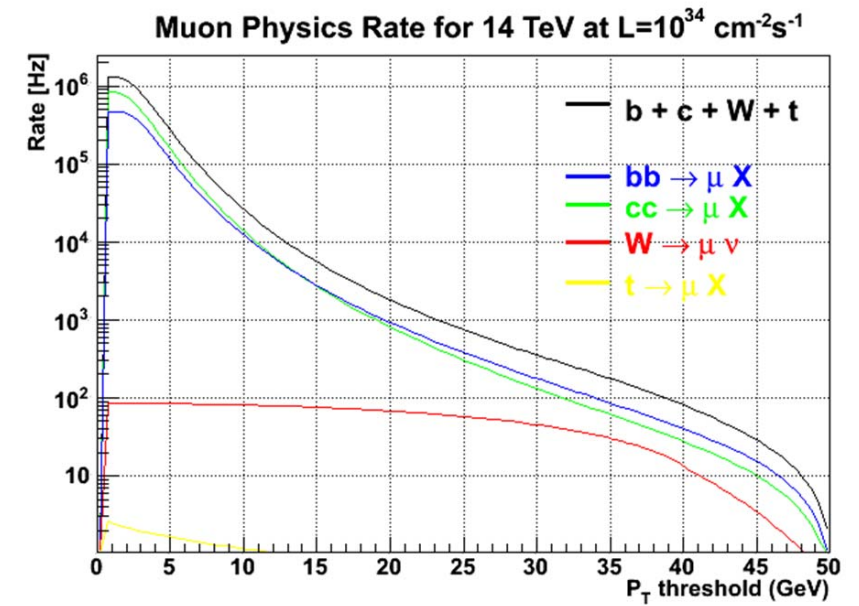
- A-C asymmetry

suggests that the background particles are charge asymmetric, combined with the toroidal field.

# Endcap muon trigger chambers



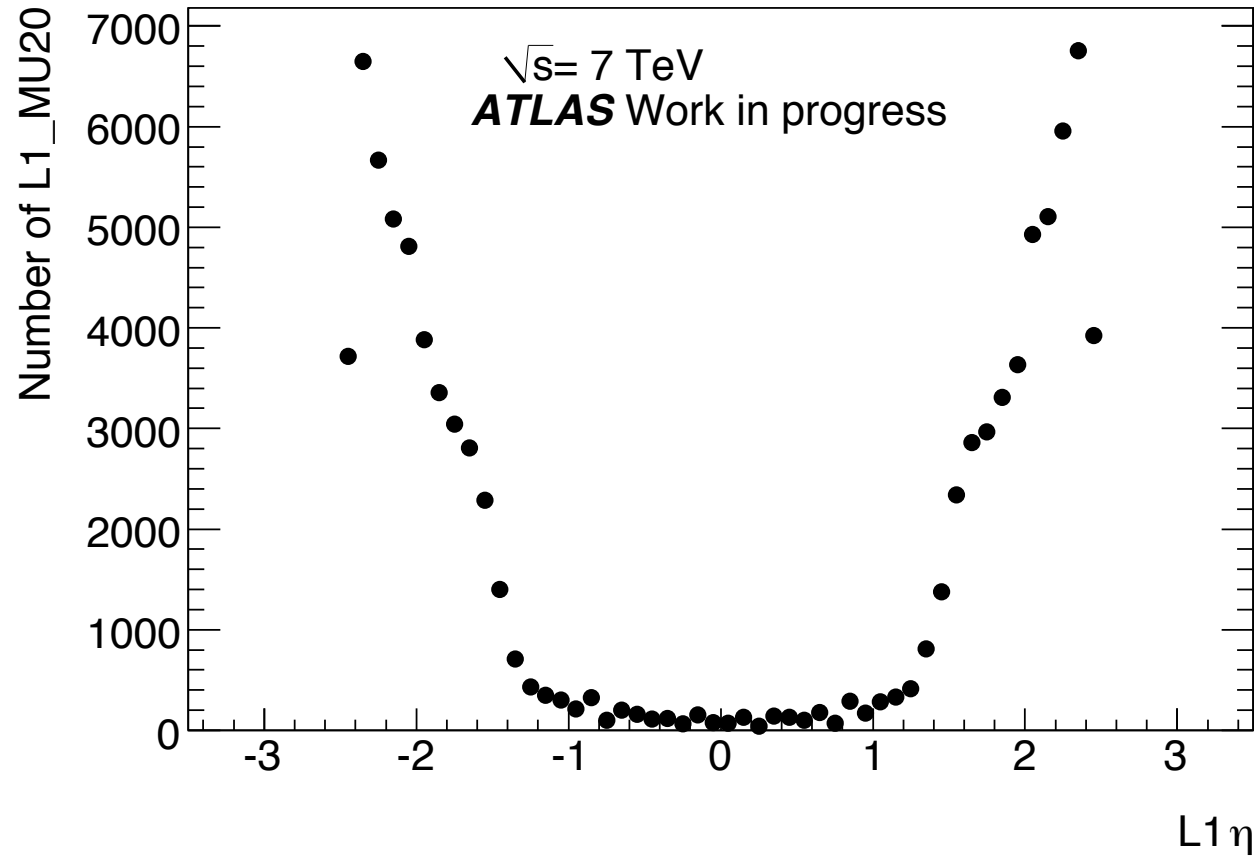
## ► Cross sections





# Muon L1

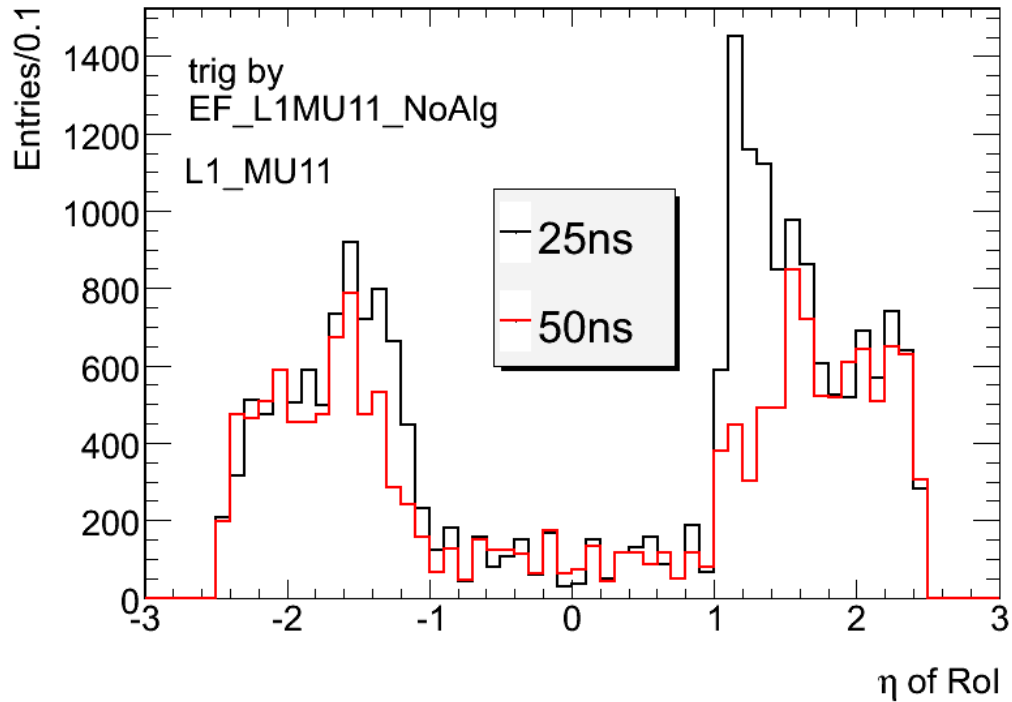
Toroidal field OFF data



F-B symmetry restores  
when magnet is OFF

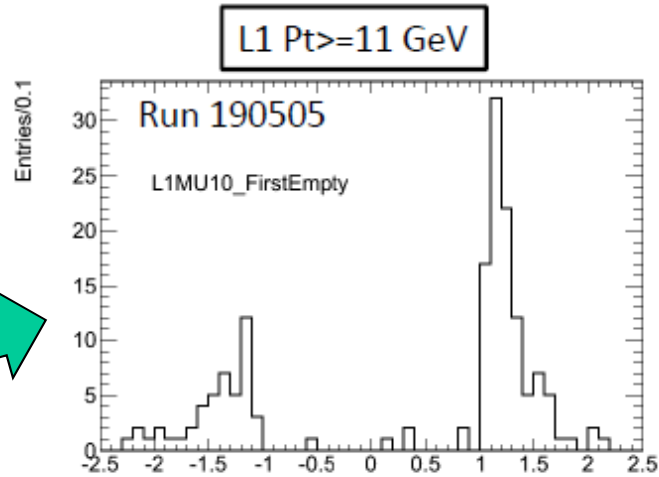
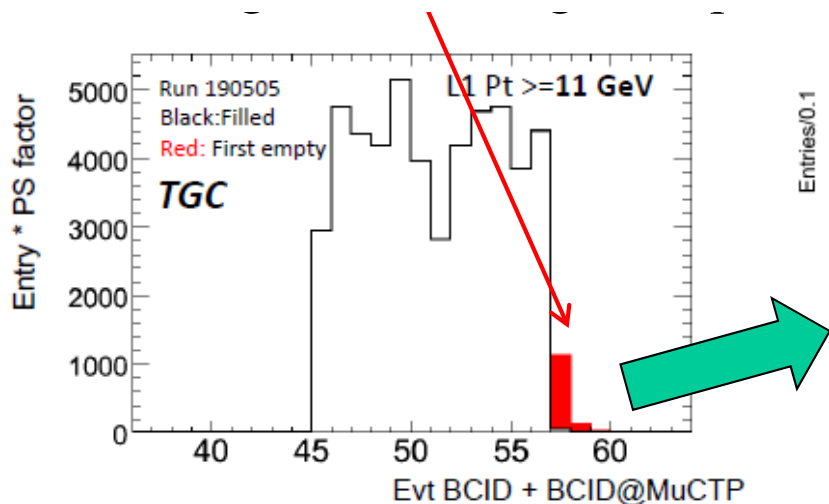
Suggests charge  
asymmetry of  
particles

# MU11 vs $\eta$ 25 ns test in 2011

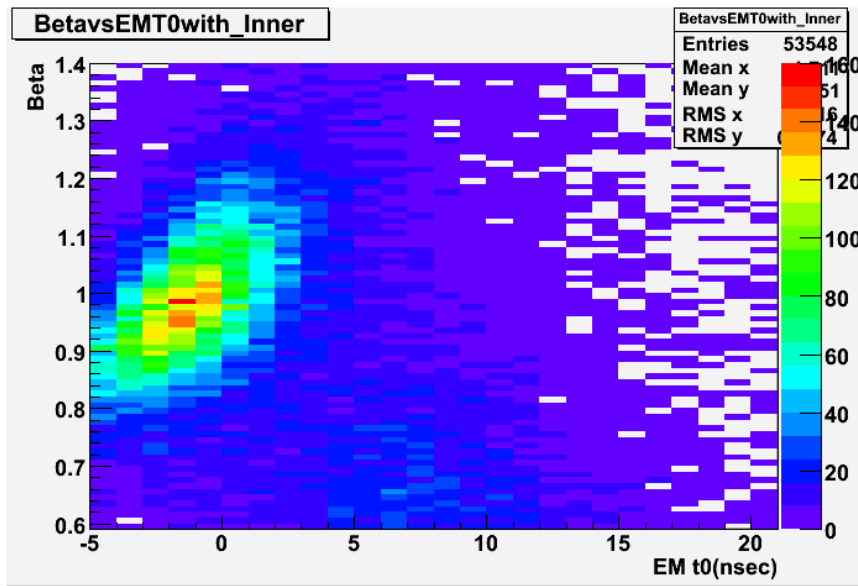


Additional rate at small  $\eta$  (large R)

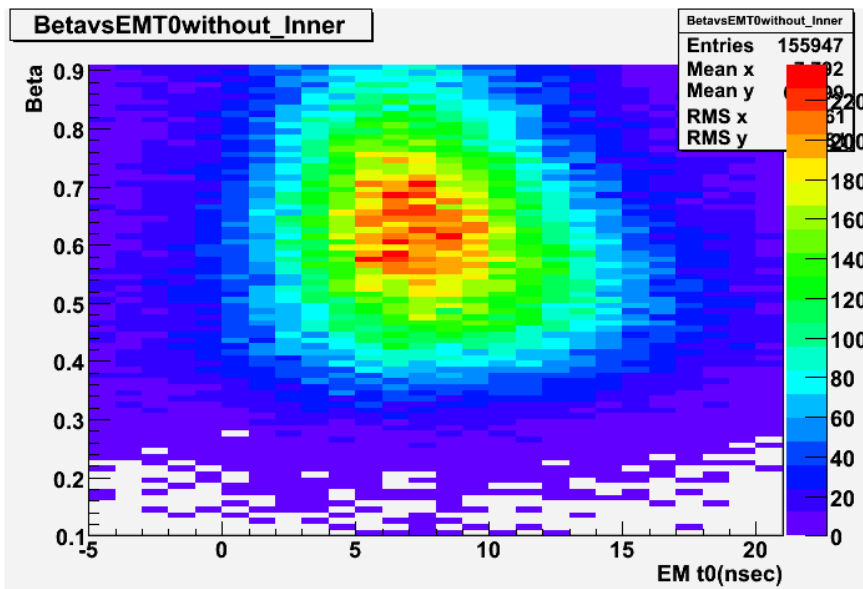
Slow particles : out of time, but captured at the next bunch crossing with 25 ns interval.



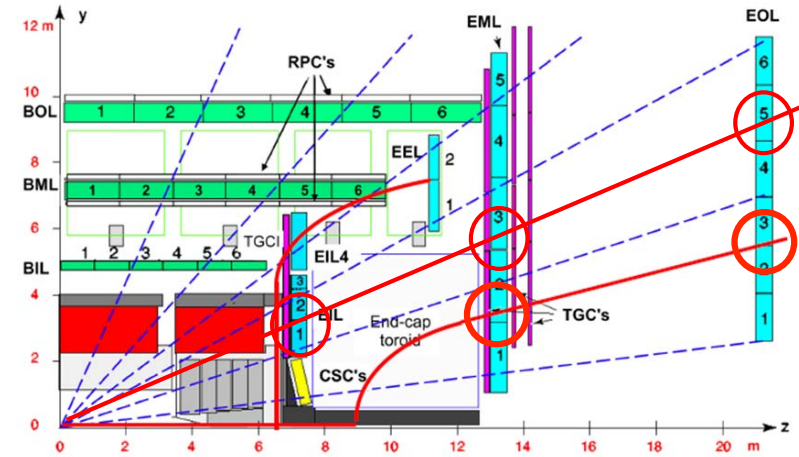
With EI hits :  $\beta \sim 1$



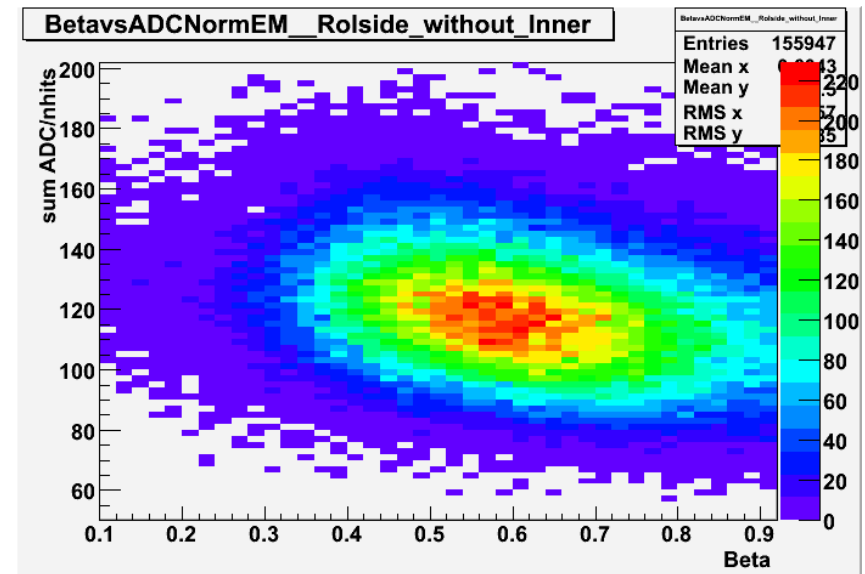
Without EI hits :  $\beta \sim 0.6$



Study of TOF between EM and EO  
(using T0 refit of MDT data)

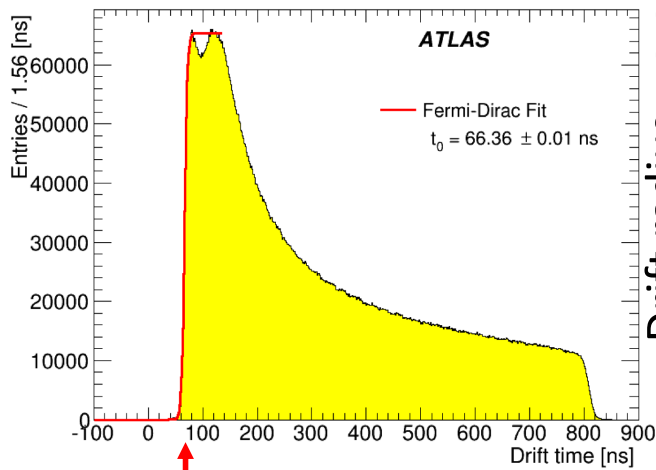
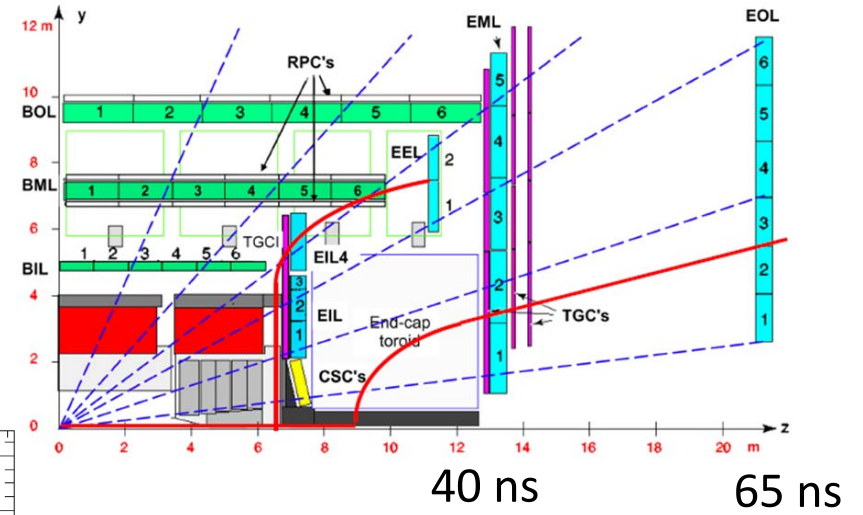
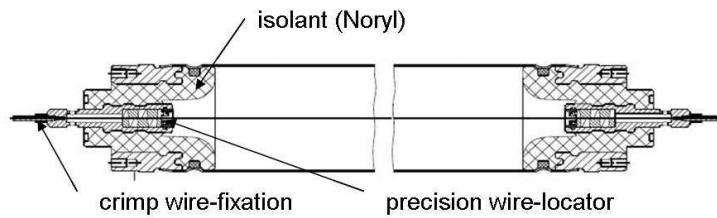
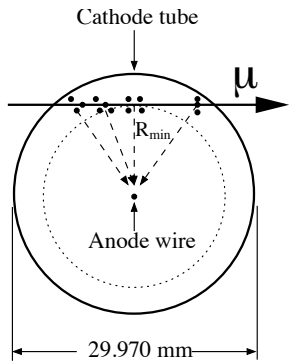


dE/dX vs  $\beta$



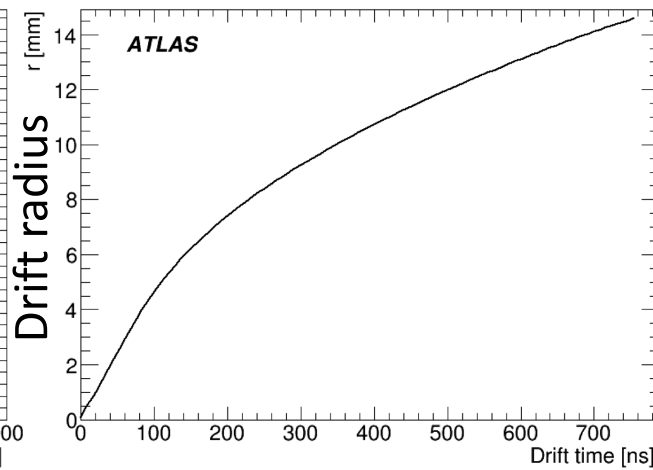
Suggests the background particles may be protons of  $p \sim 1$  GeV



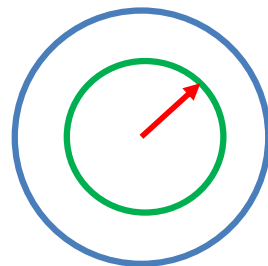
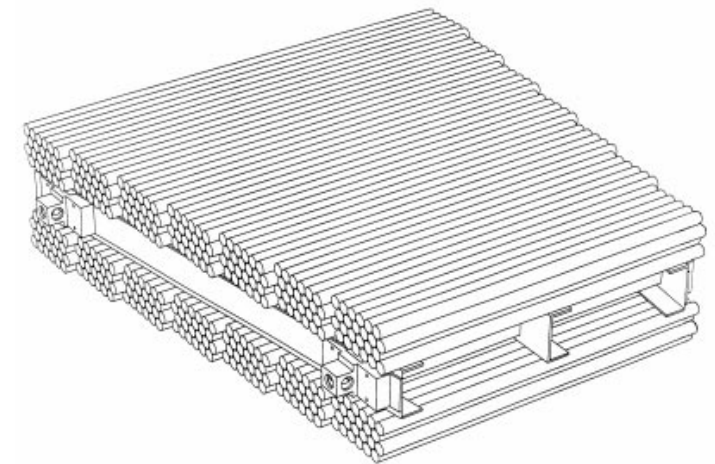


time spectrum

$t_0$



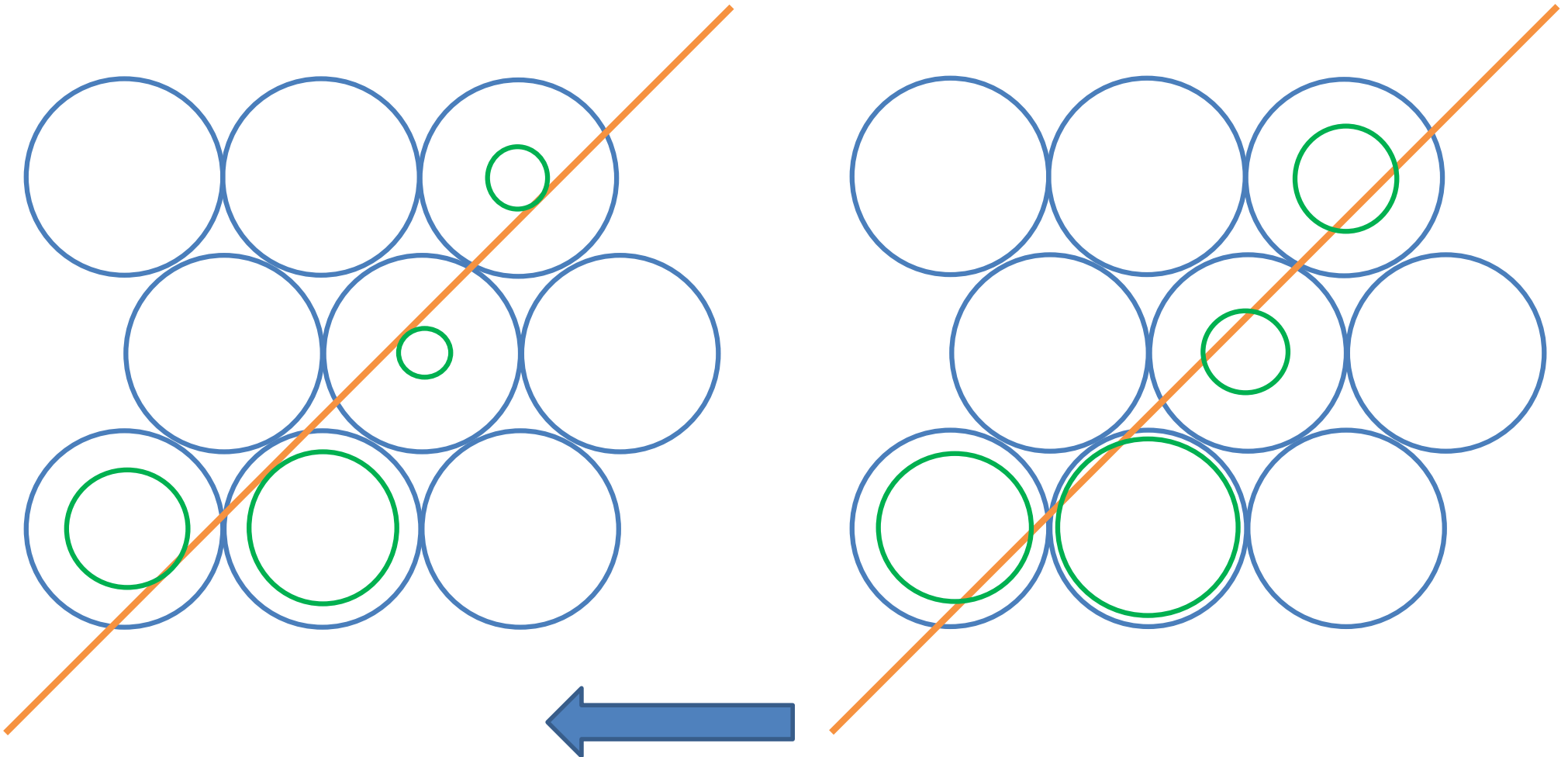
Drift time



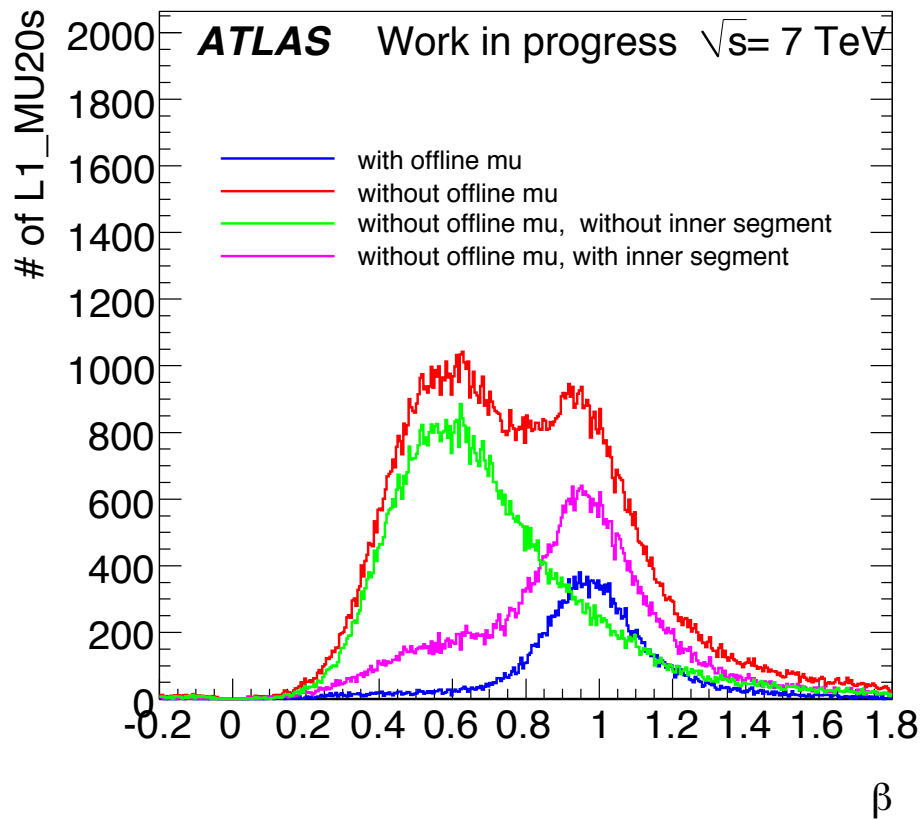
Time measurement  $\rightarrow$  Drift circle

A particle of correct timing ( $\beta = 1$ )

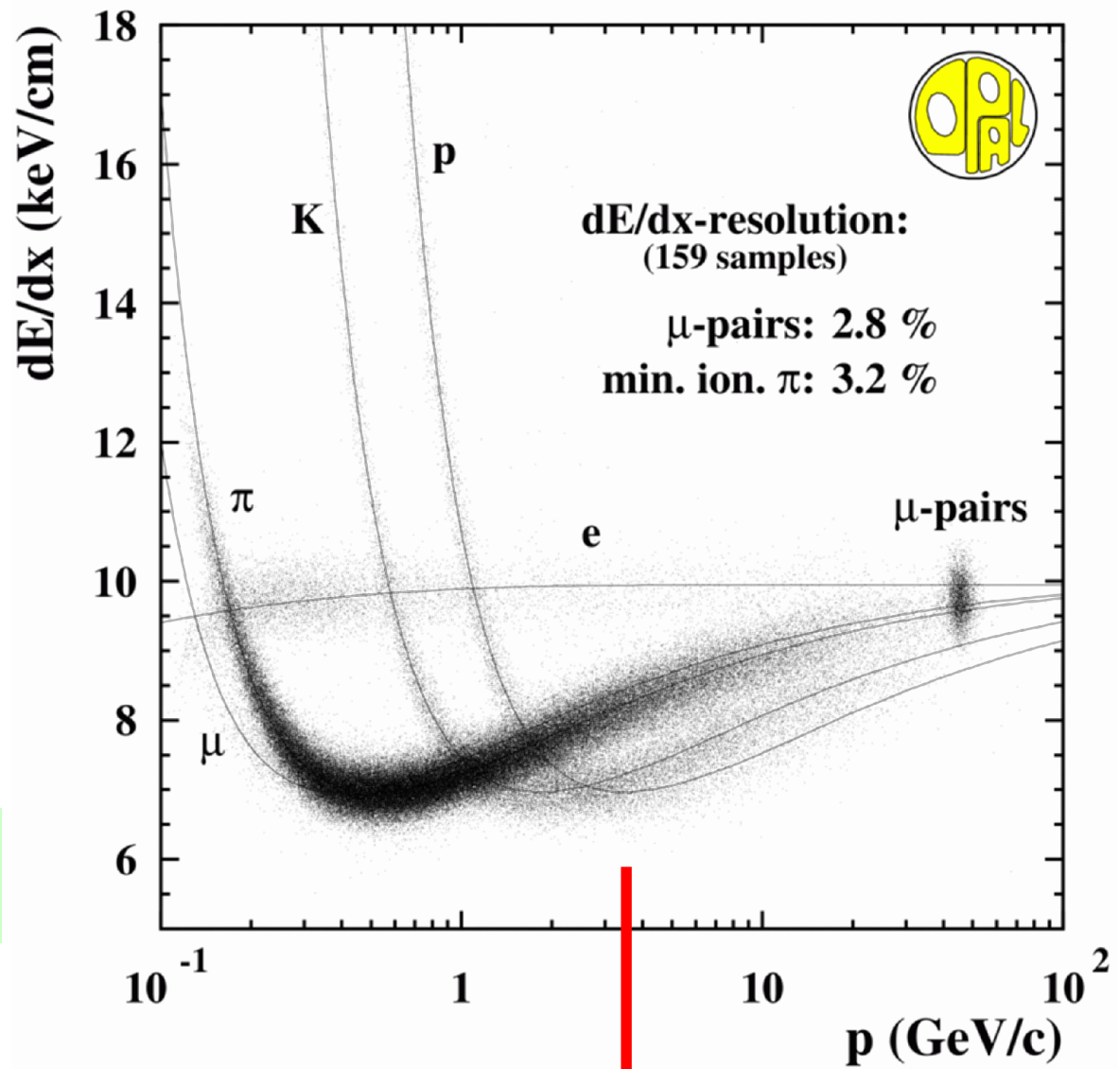
Late arriving particle ( $\beta < 1$ )



$t_0$  adjustment ( $t_0$  refit) : determination of precise timing

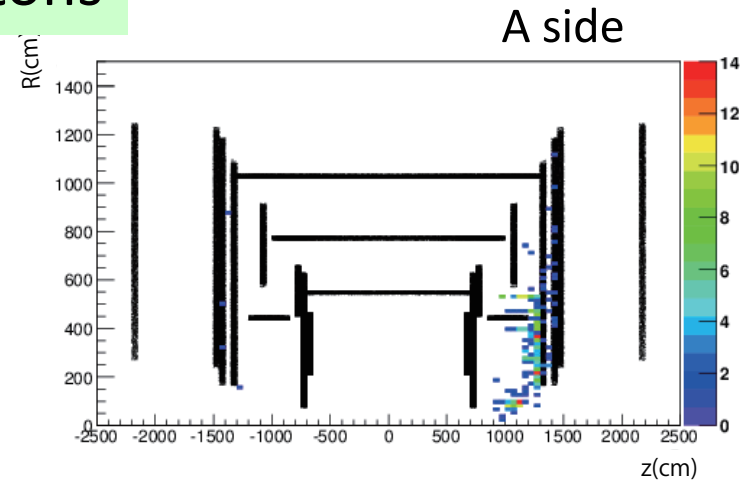
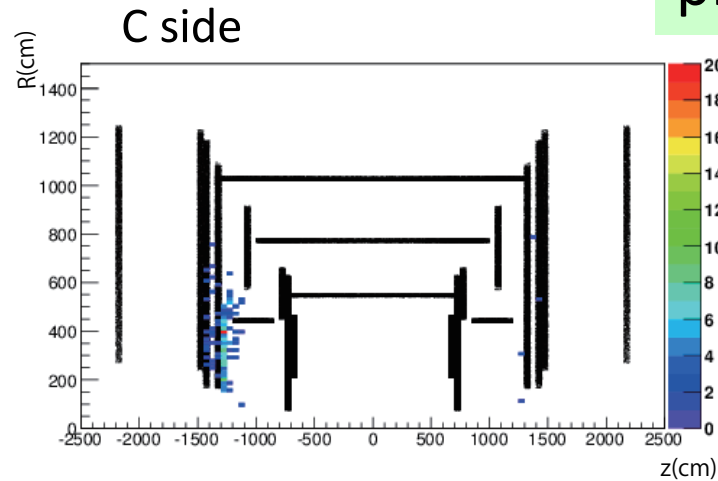


- Without offline  $\mu$ , without EI hits
- With offline  $\mu$ , with EI hits



# Monte Carlo simulation

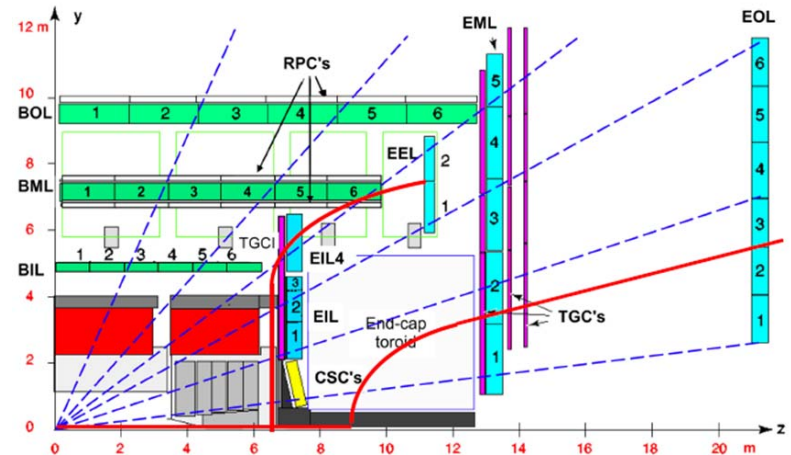
protons



Map of birth positions of fake trigger particles

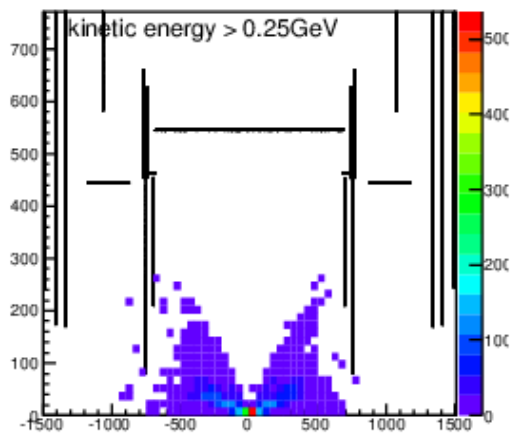
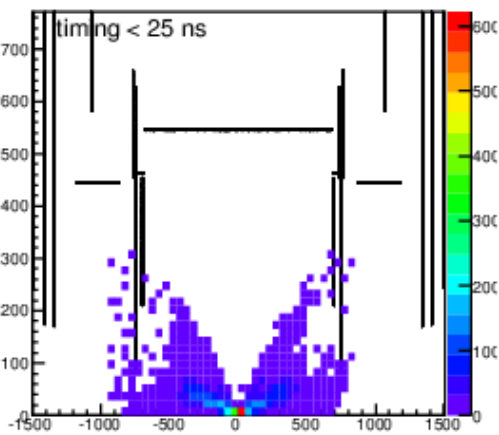
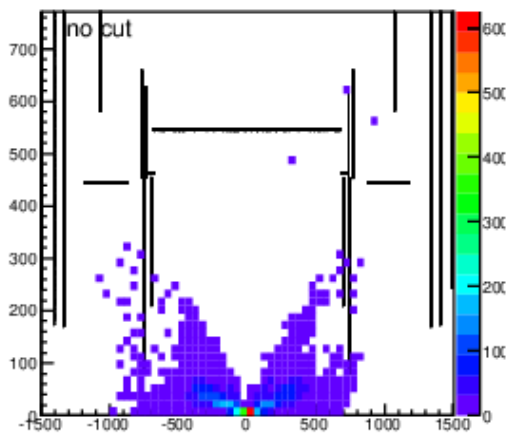
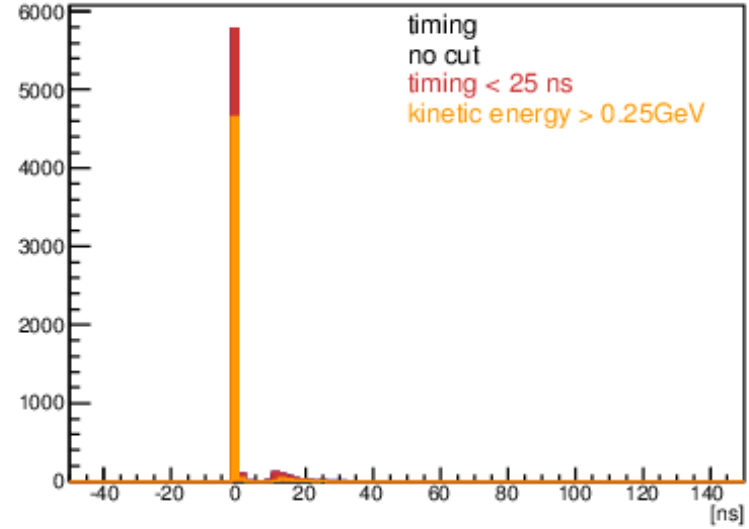
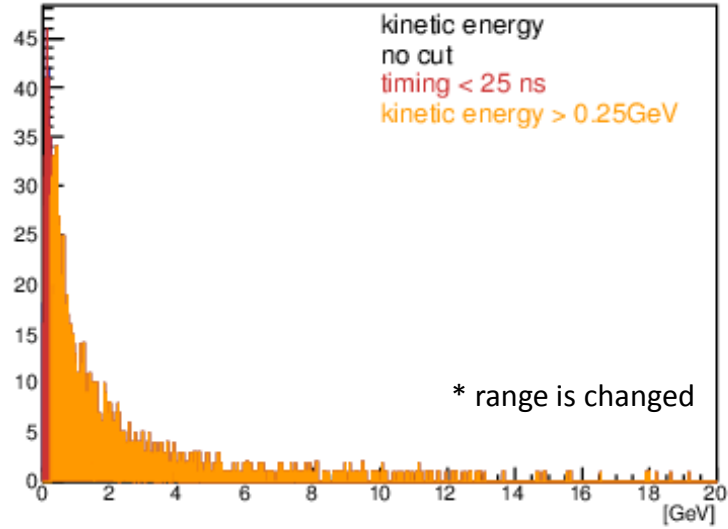
## Interpretation

Fake muons are due to protons resulting from complex nuclear reactions in front of EM station.

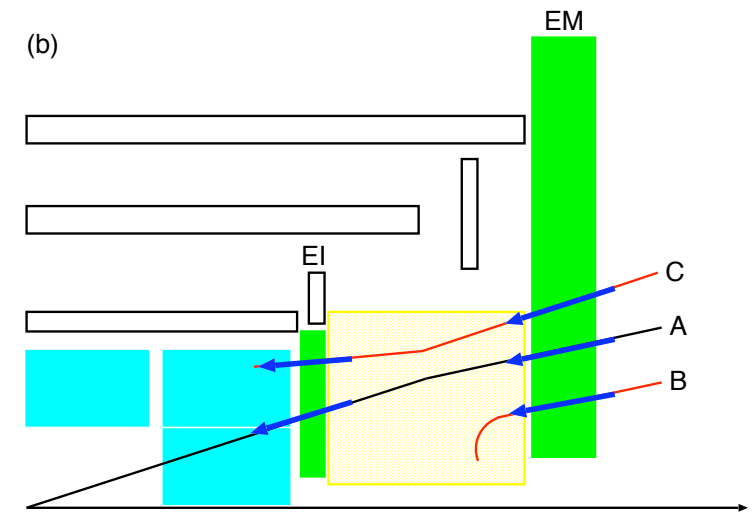




for comparison,  $\mu^+$



Doing these studies and considerations, idea converged to a new small wheels.



- Precision tracker that works at high rate (including the phase 2 environment)
- Provide real time EI segments to validate the EM L1 segments
- Tracking performance as good as the existing small wheel

+ good online angular resolution (1 mrad) to allow better  $P_T$  resolution (sharper L1 threshold) at Phase-2 combined with improved EM angular resolution (e.g. using MDT).

Proposals of detector technologies and launch of R&Ds

sMDT(small tube), sTGC(small strip), mRPC(narrow gap), MM(mpgd), ....

## Several detector options

- sMDT (tracking) + sTGC (trigger + tracking  $\phi$  &  $R$ )
- sMDT (tracking) + mRPC (trigger + tracking  $\phi$ )
- MM (tracking + trigger)

## Towards TDR, technology choice

Workshop at Le Brassus (Suisse Jura) Jan. 2012

→ instead of converging, .. another proposal sMDT+sTGC (Outer) + MM (Inner)

Totally confused. → Further discussions

Eventually converged to the present NSW design after ~ 1 year of many discussions and additional R&D milestones

High redundancy was one of the reasons.

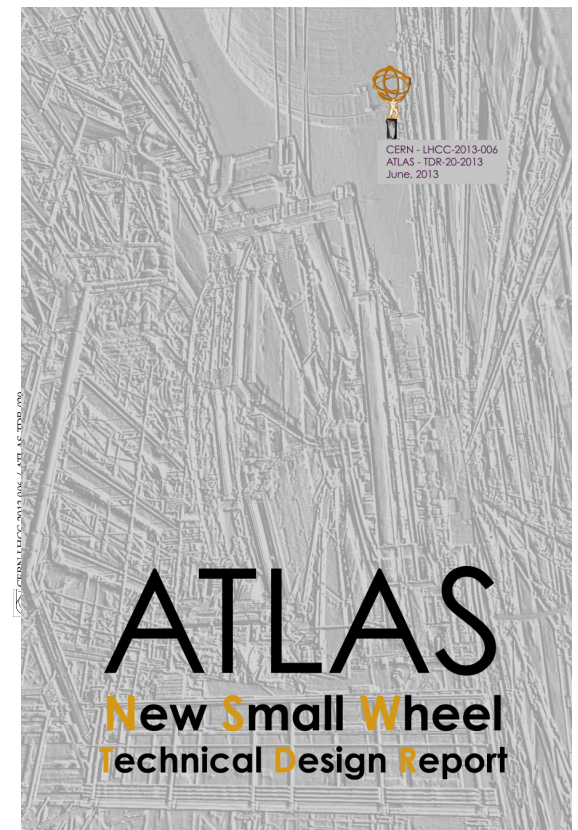
two technologies for both tracking + trigger

# Technical Design Report

- Initial design review : with sTGC + MM proposal, but with milestones to be achieved by the end of the year
- Kick off institute meeting August 2012 ~ 50 institutes
- NSW milestone review Jan 2013 Aix les Bains
- EB approval
- CB approval
- TDR & iMoU

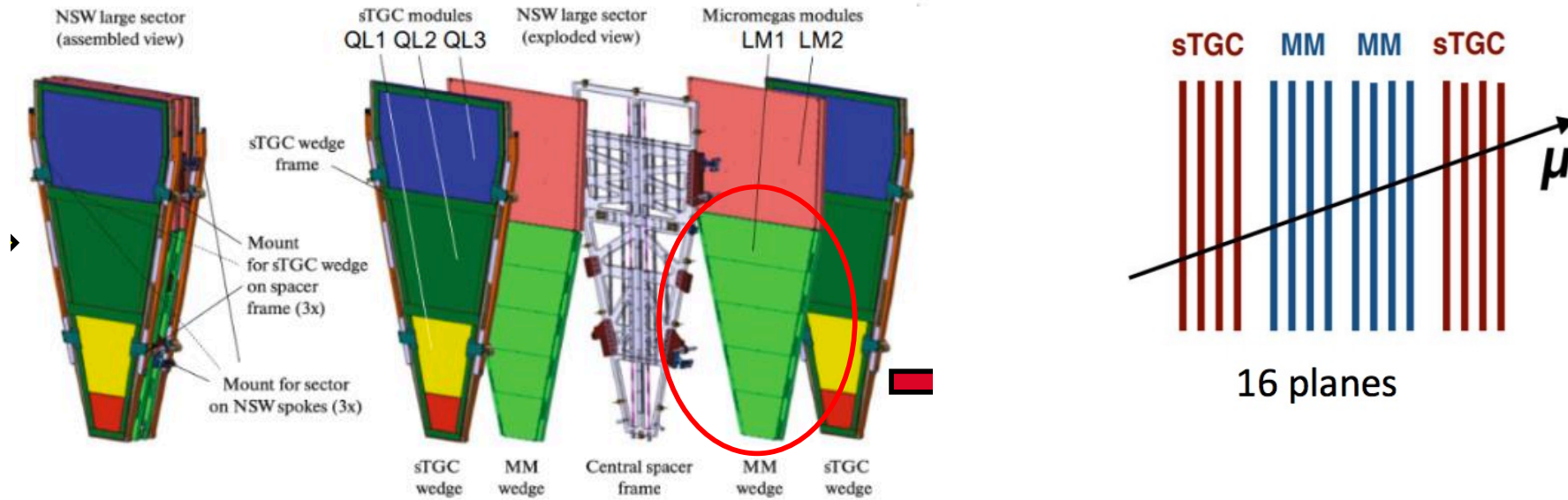
NSW TDR CERN-LHCC-2013-006, June 2013

In the meantime, the TGC of the old SW and Tile signal were included in L1 in coincidence with EM → reduction of rate as expected, Justifying the basic idea of NSW for rate reduction.



# NSW overview

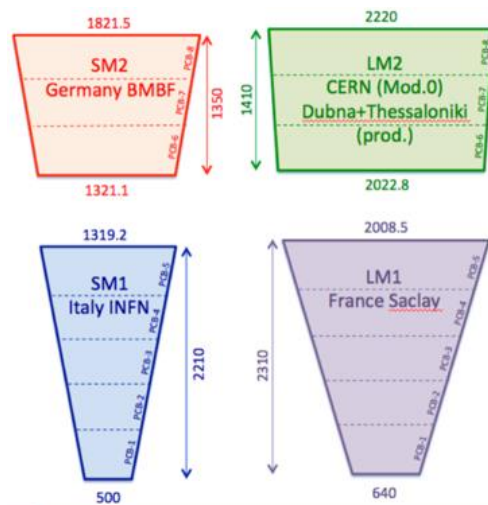
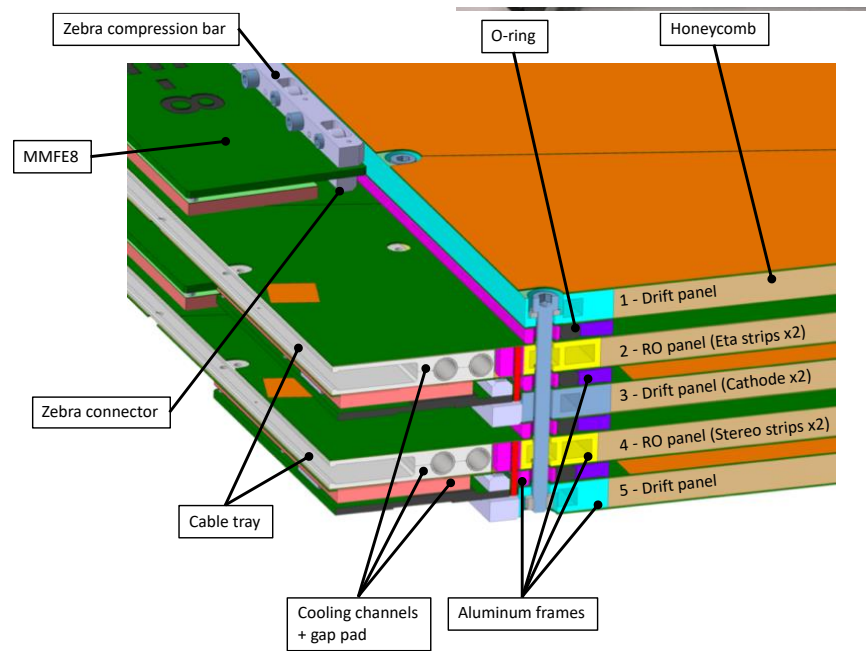
I myself was involved in development and construction of MM chambers at Tokyo and Saclay





# MM chambers

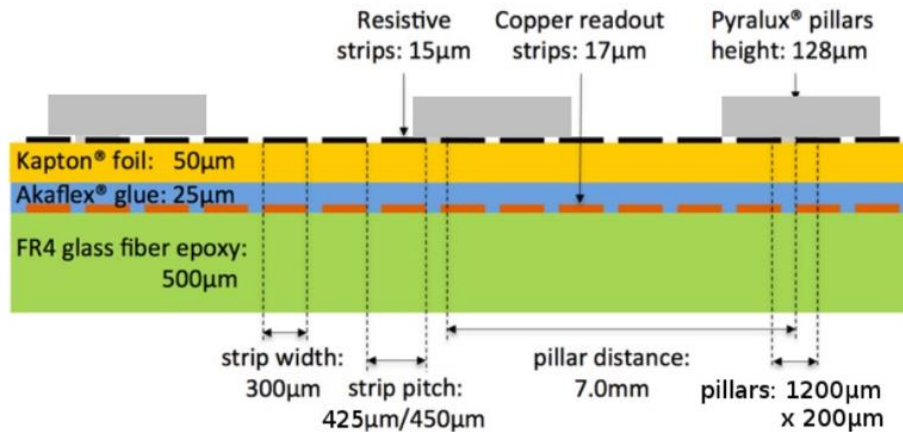
## Basic quadruplet unit



### 4 construction sites:

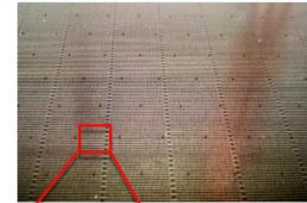
- SM1 → Italy/INFN
- SM2 → Germany
- LM1 → France/Saclay
- LM2 → Russia/Dubna – Greece/Thessaloniki (+ Cern)

# MM chambers



- 1022 strips/board
- Readout strips: 300 µm width, 425 or 450µm pitch
- Screen-printed resistive strip pattern with same pitch
- Resistive strips interconnected; pattern interrupted in the center → two HV sections per board
- HV supply via silver line from the side
- Elongated pillars: 1200µm x 200µm
- Readout strips routed to pads for elastomeric connectors (Zebra)

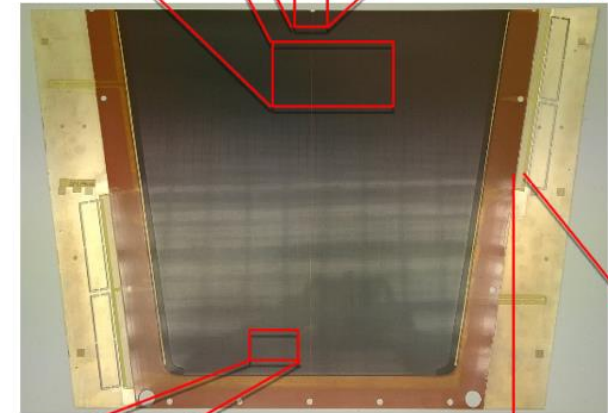
Surface structure of the board, with regular patterns of pillars and interconnections of resistive strips



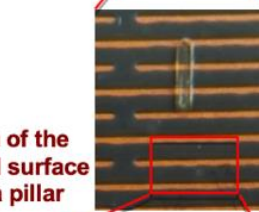
Central upper hole for layer interconnection. Readout strips routed around the hole. Central separation of resistive pattern.



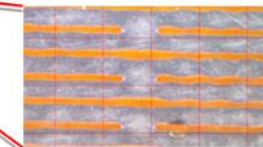
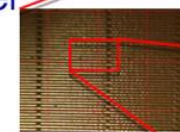
Example of NSW Micromegas readout board (smallest type)



Zoom of the board surface with a pillar



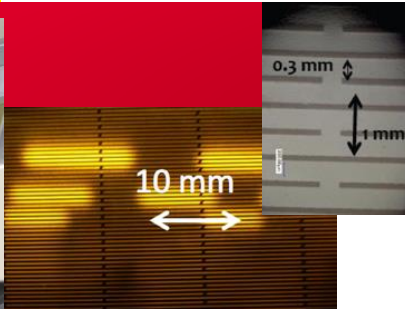
Zoom of the resistive strip interconnections



Pads for elastomeric connector (Zebra)



# MM chambers



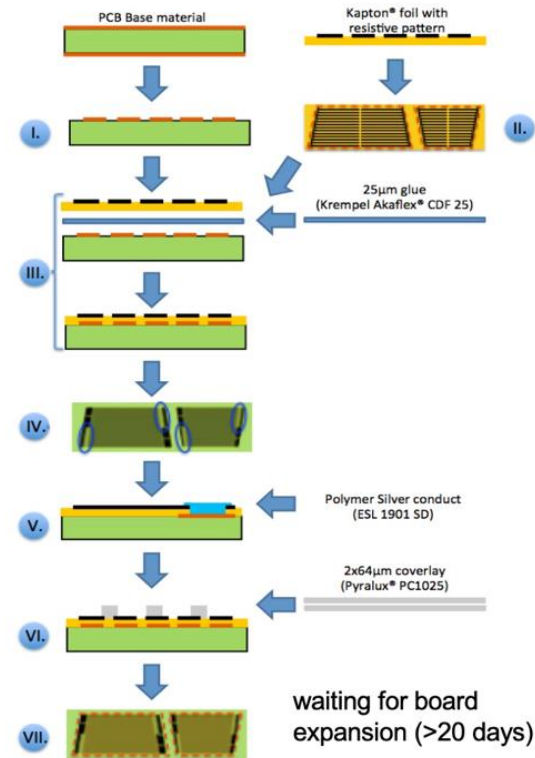
# BOARD PRODUCTION PROCESS



Typical resistivity ~ 10-20 MΩ/cm (~800 kΩ/□)

- I. Photolithographic creation of copper pattern  
standard process.  
complex due to: size of board, required precision & board elongation (humidity).
- II. Cutting of Kapton foil with resistive pattern  
non-standard but simple & required accuracy only ±1mm
- III. Stacking and gluing at high P&T of Kapton foil, glue foil and board  
standard process for small boards  
complex due to: size of board & required cleanliness.
- IV. Chemical silver plating of copper pads  
standard process
- V. Screen-printing of silver paste  
non-standard but rather simple & required accuracy only ± 1mm
- VI. Lamination of coverlay & pillar creation  
standard process for small boards.  
complex due to: size of boards, highly non-standard pattern, required flatness
- VII. Cutting of boards and drilling of non-precision holes  
standard process on CNC machine.  
complex due to size of boards, required cutting precision & board elongation (humidity).

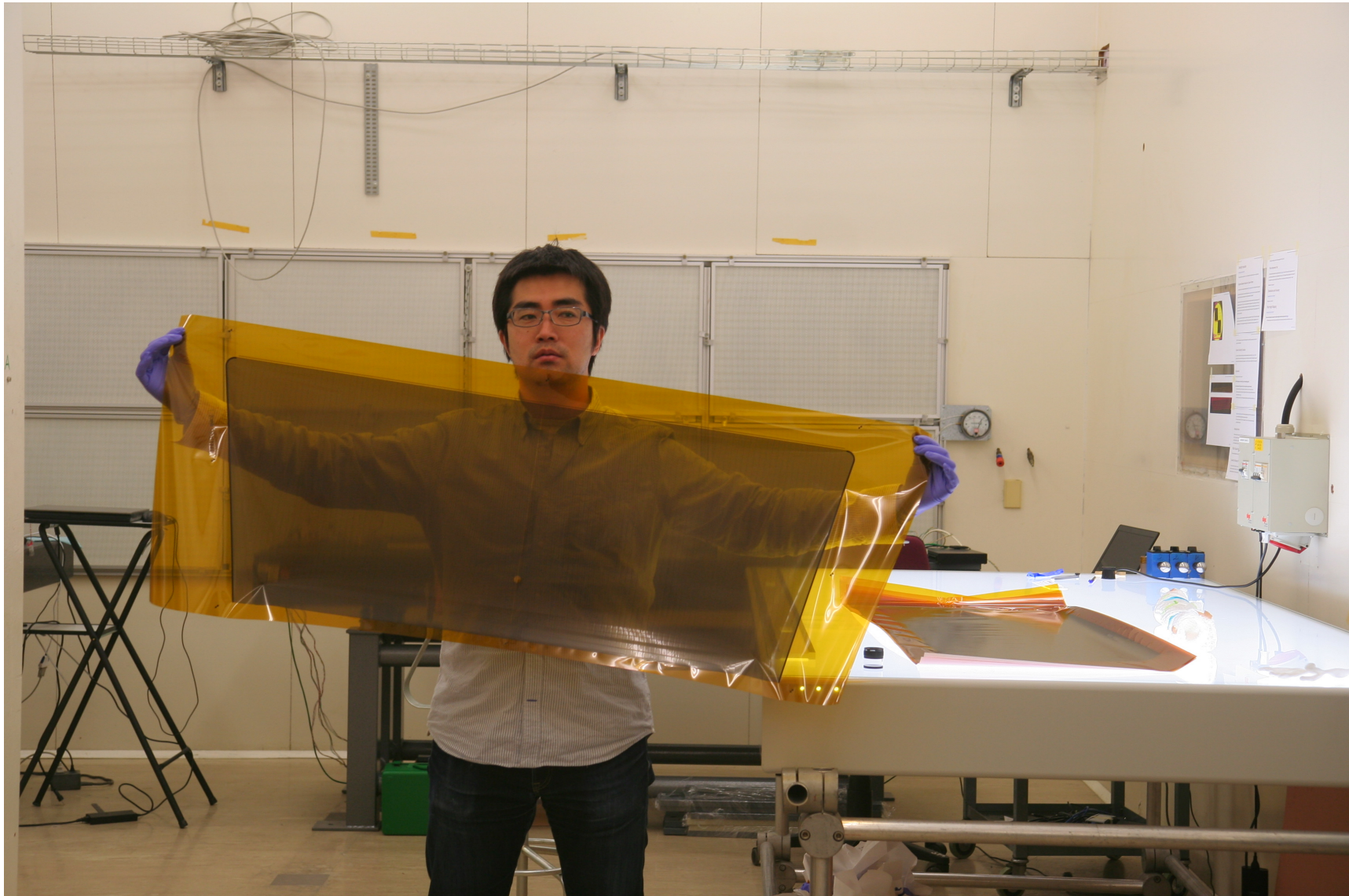
Try to make all production steps as close as possible to standard processes in industry





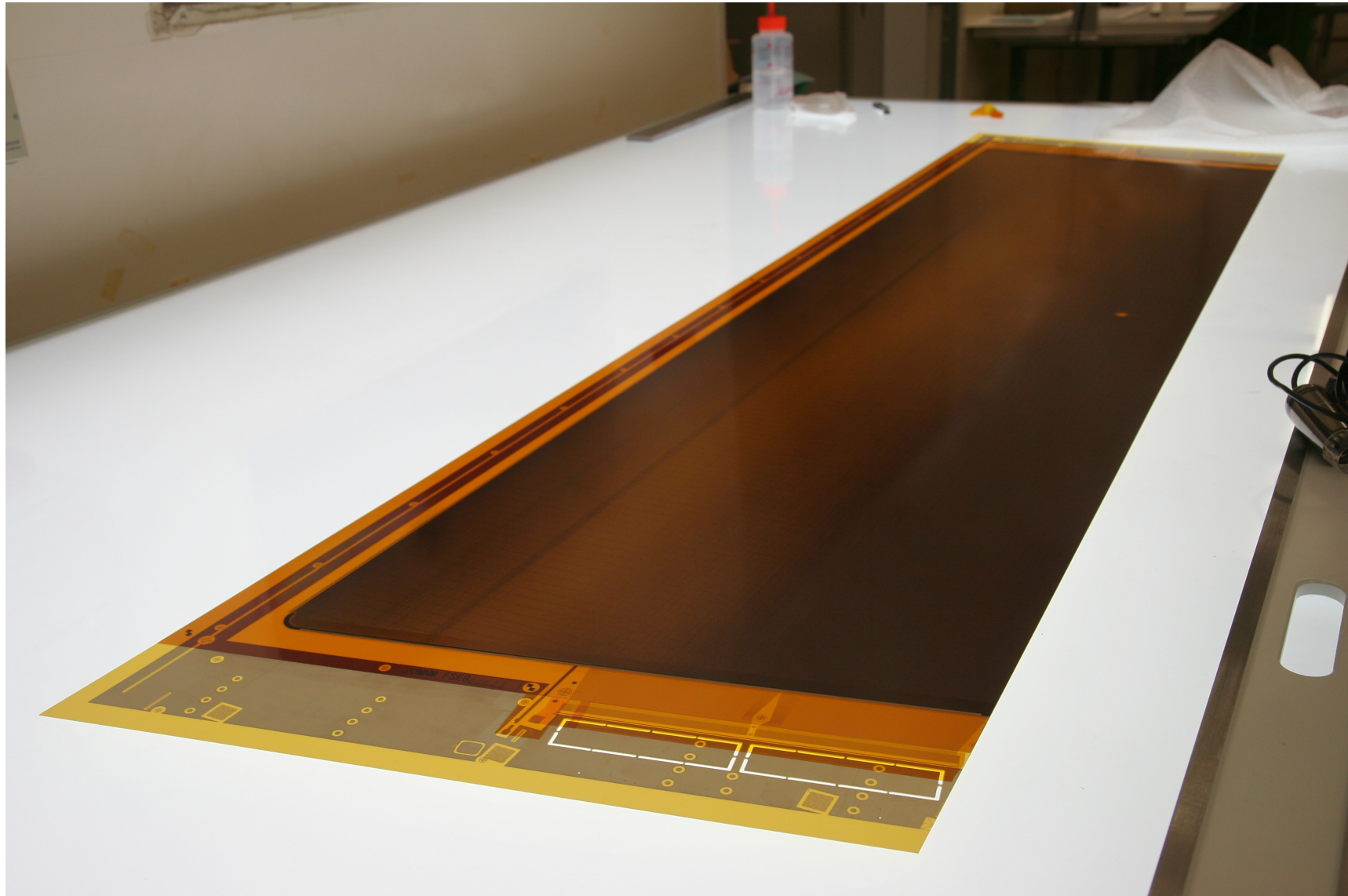
# MM chambers

Prof. Masubuchi, U.Tokyo



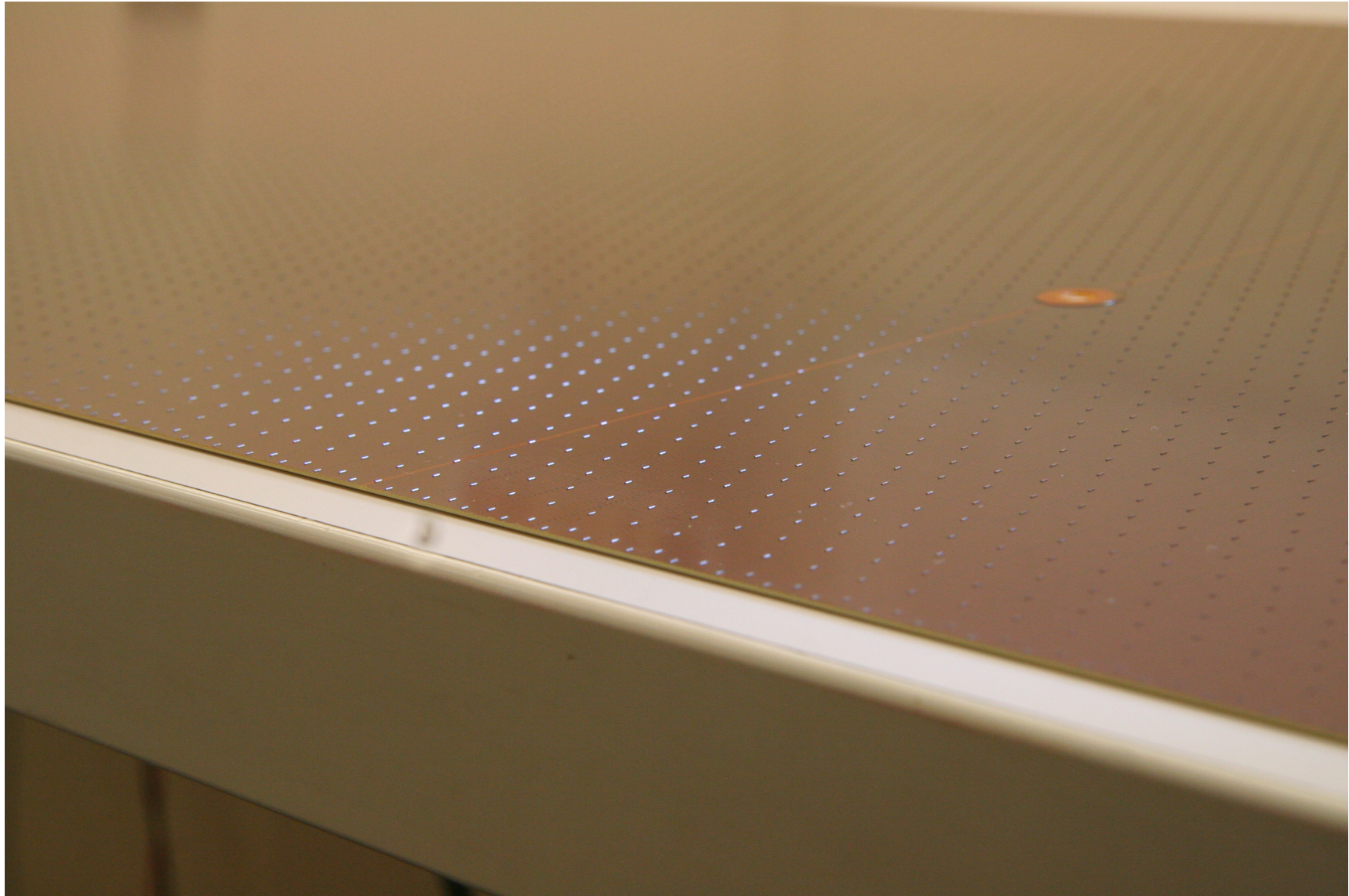


# MM chambers





# MM chambers

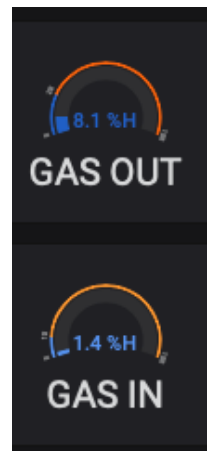
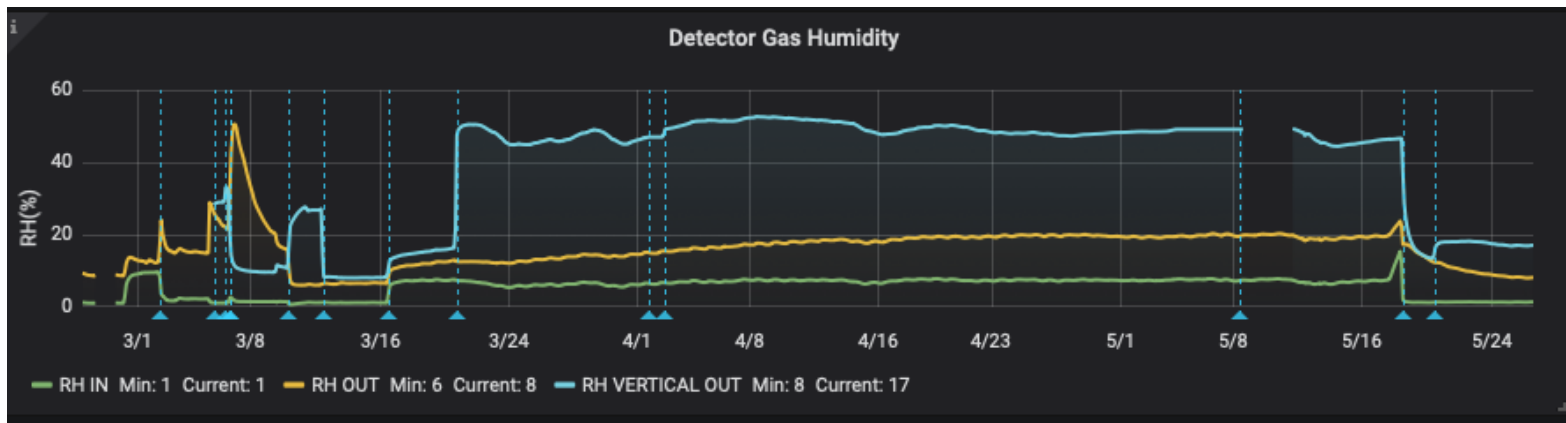


# MM HV issue

## HV INSTABILITIES: CAUSES



- ❑ Residual ionic contamination → cleaning procedures reviewed
- ❑ Mesh mechanical imperfections → mesh polishing
- ❑ Humidity → monitor humidity, dry panels and modules, increase gas flux
- ❑ Low resistivity of anode resistive strips
- ❑ ~~Low quenching gas mixture~~

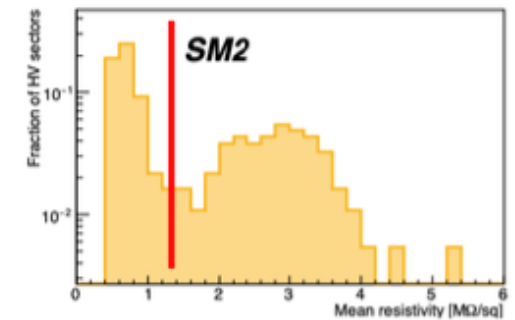
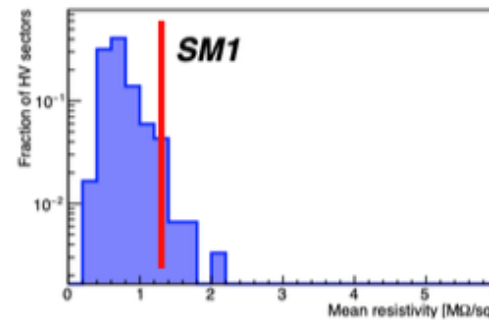
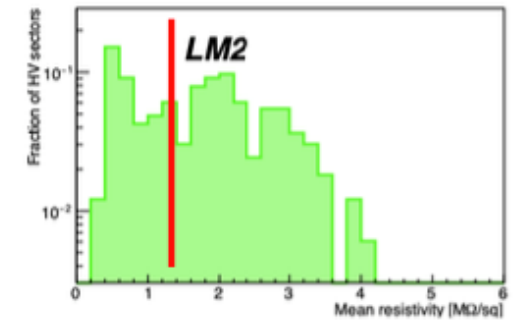
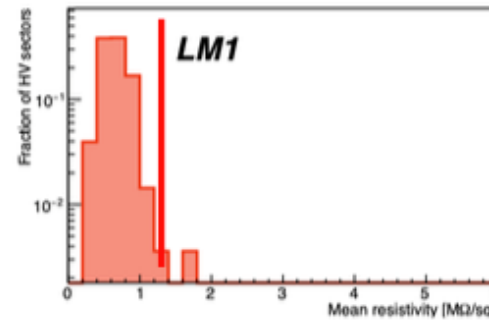
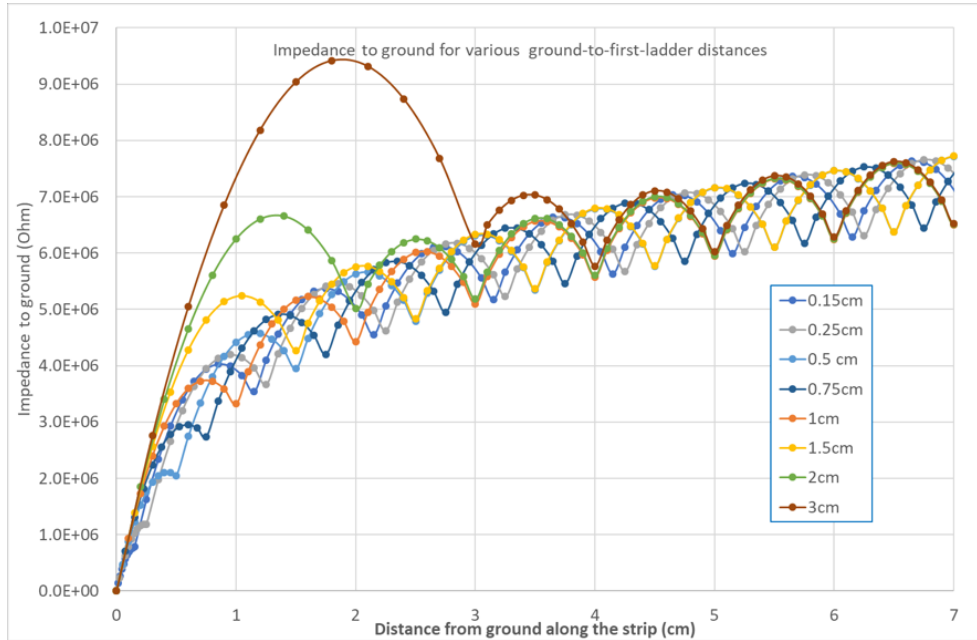
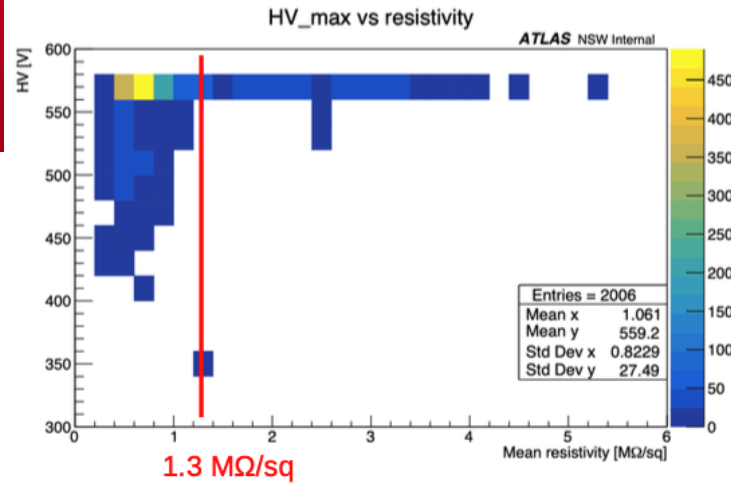
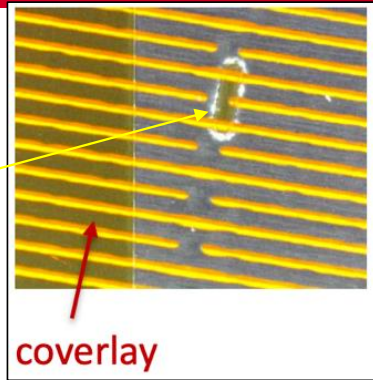




# MM HV issue



## HV INSTABILITIES

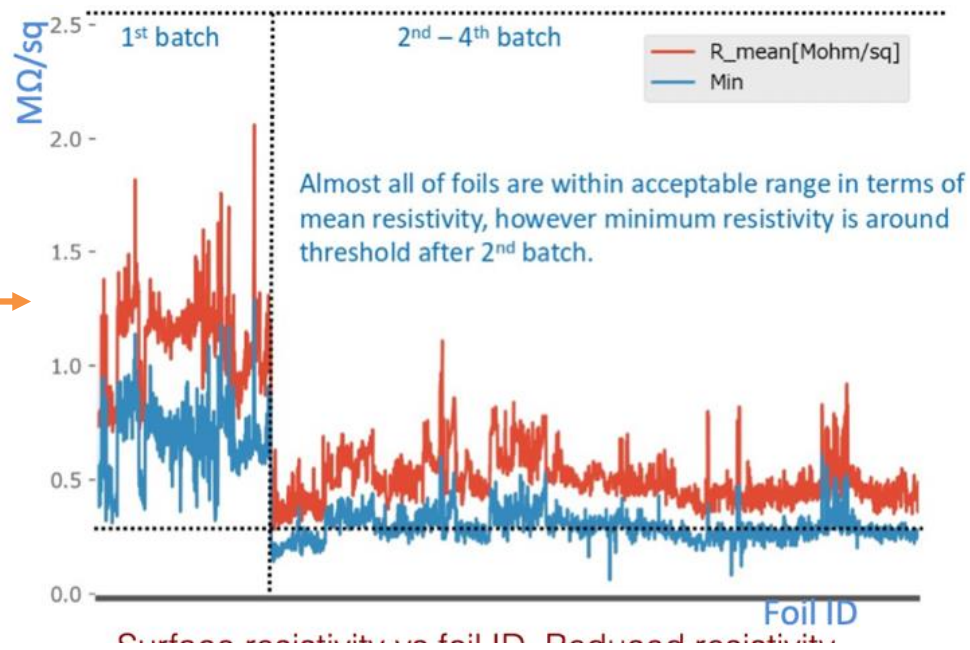


# MM HV issue

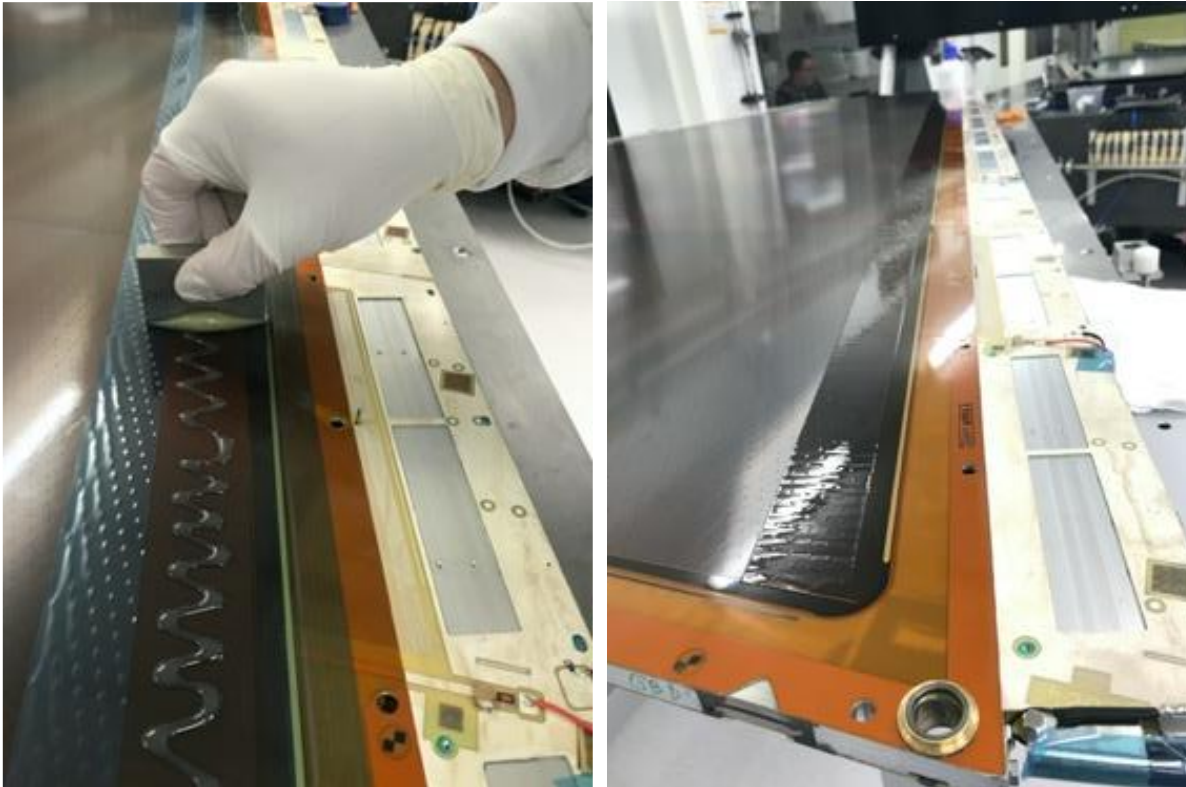
It was difficult to control the resistivity  
Batch to batch variation of resistive paste.

+

Initially defined tolerance was not quite correct



## MM HV issue

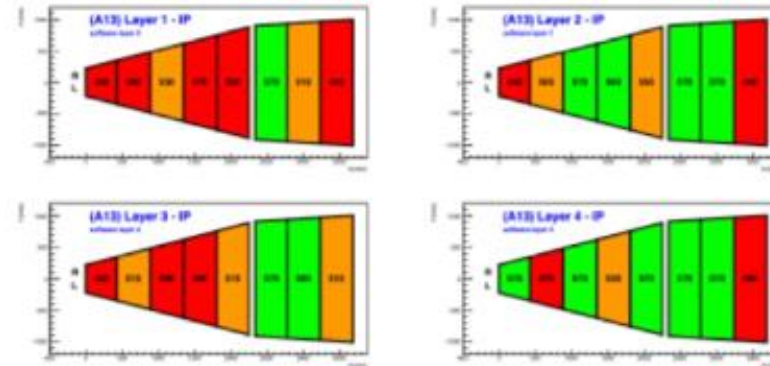


Passivation  
To ensure sufficient  
resistance from the silver  
line,  
sacrificing the active area.



# MM new gas mixture to save

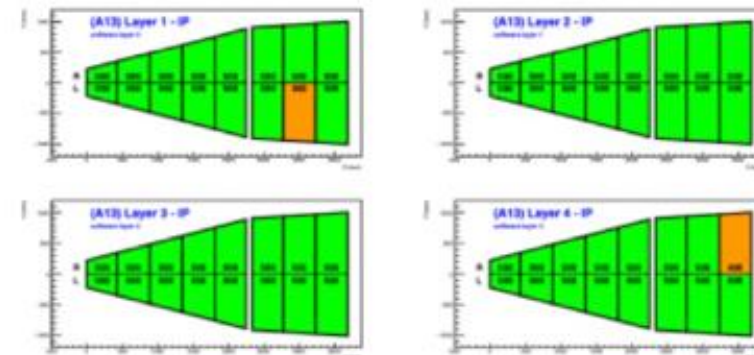
Ar:CO<sub>2</sub> 93:7



Adding 2% of iC<sub>4</sub>H<sub>10</sub> greatly improved the HV stability + same gain at lower HV.

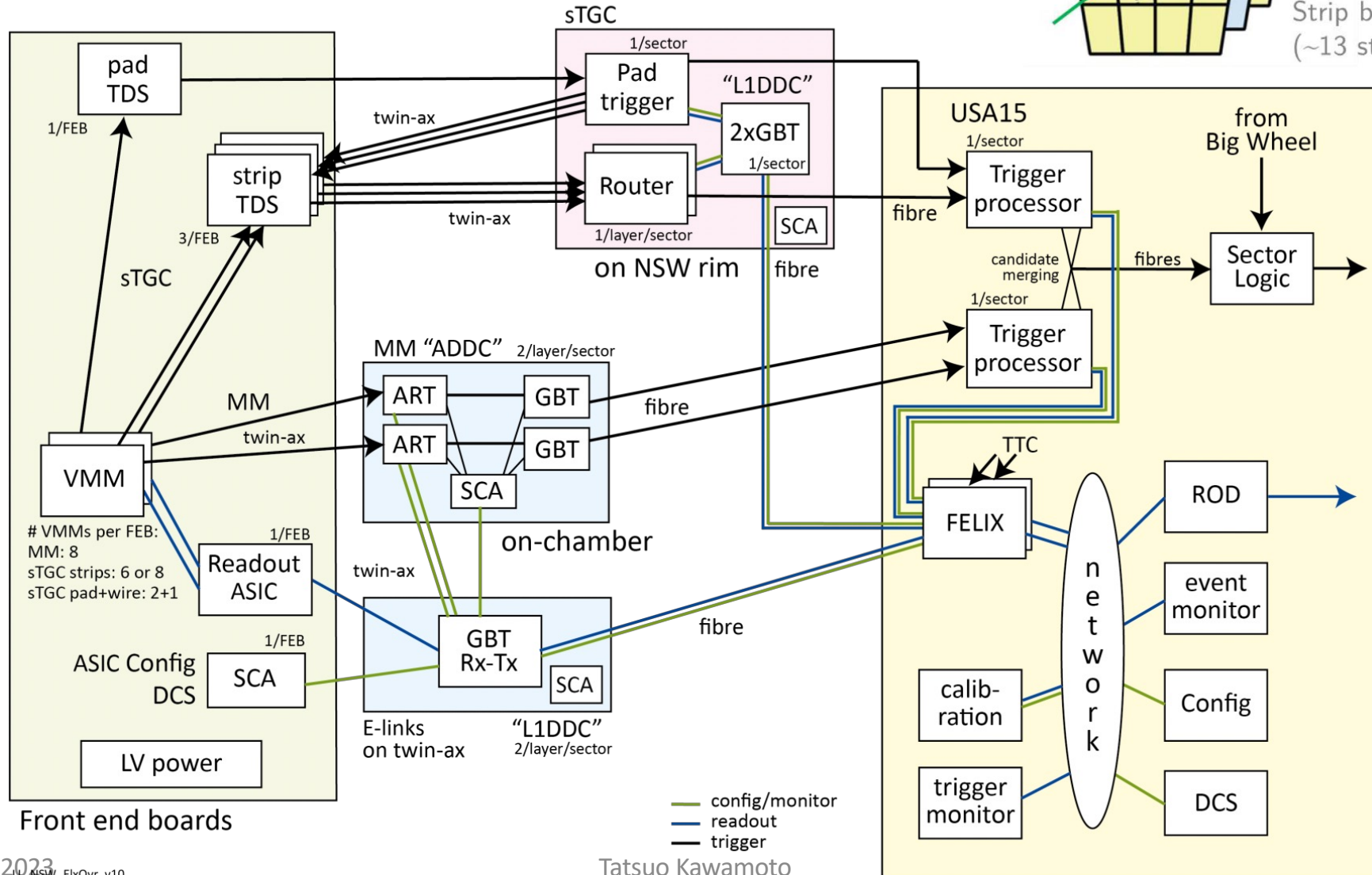
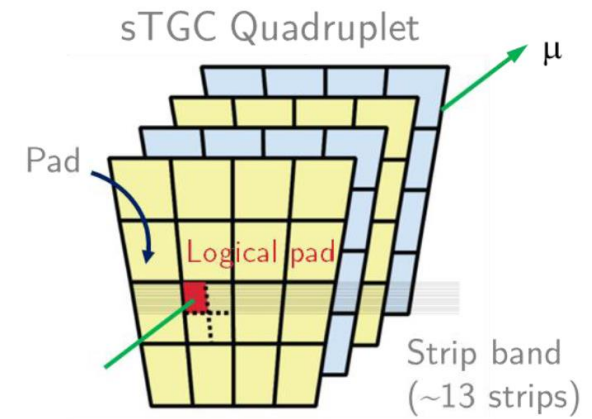
Took ~1 year to convince people and ourself to switch the gas mixture. with Irradiation tests, etc.

Ar:CO<sub>2</sub>iC<sub>4</sub>H<sub>10</sub> 93:5:2



# NSW electronics

## NSW Electronics Trigger & DAQ dataflow



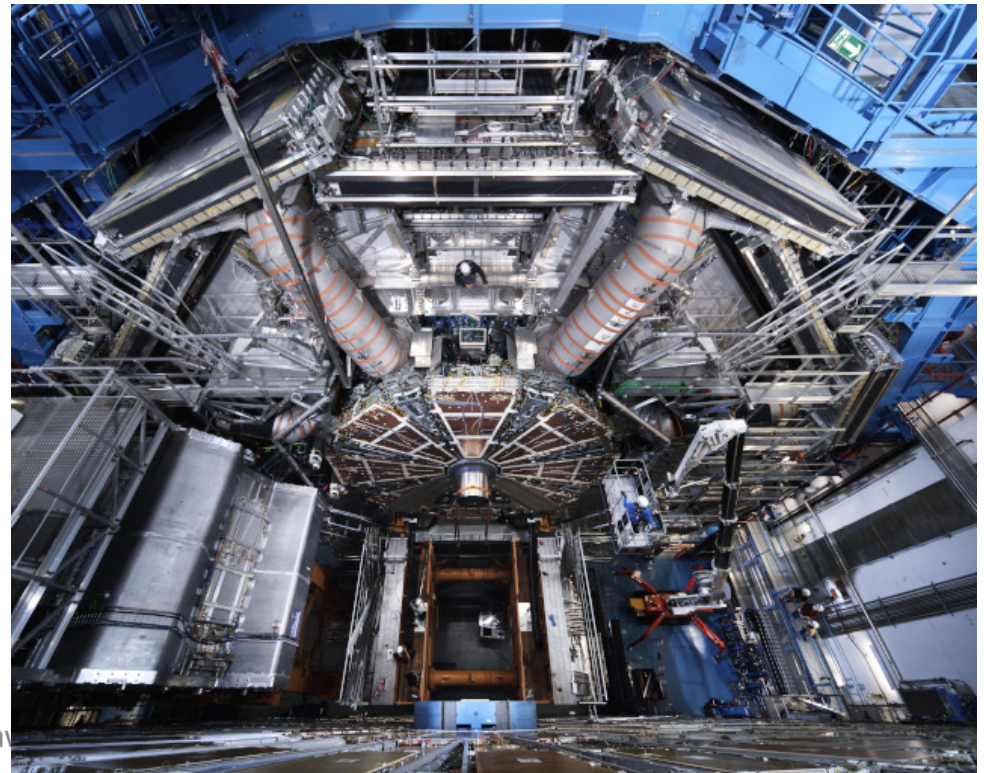
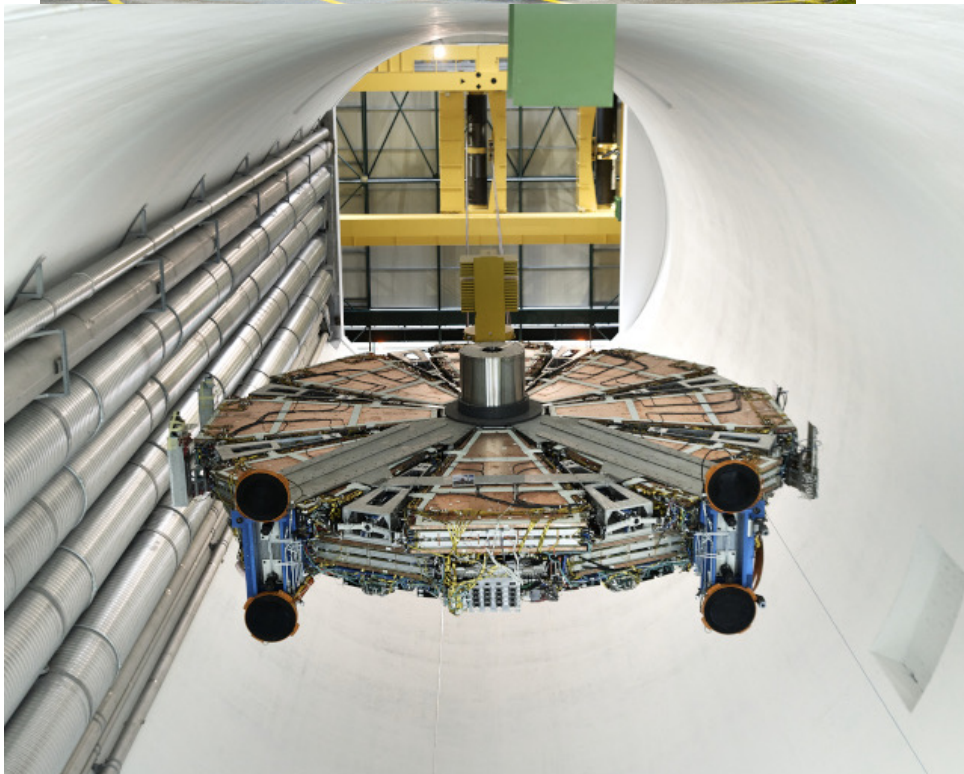
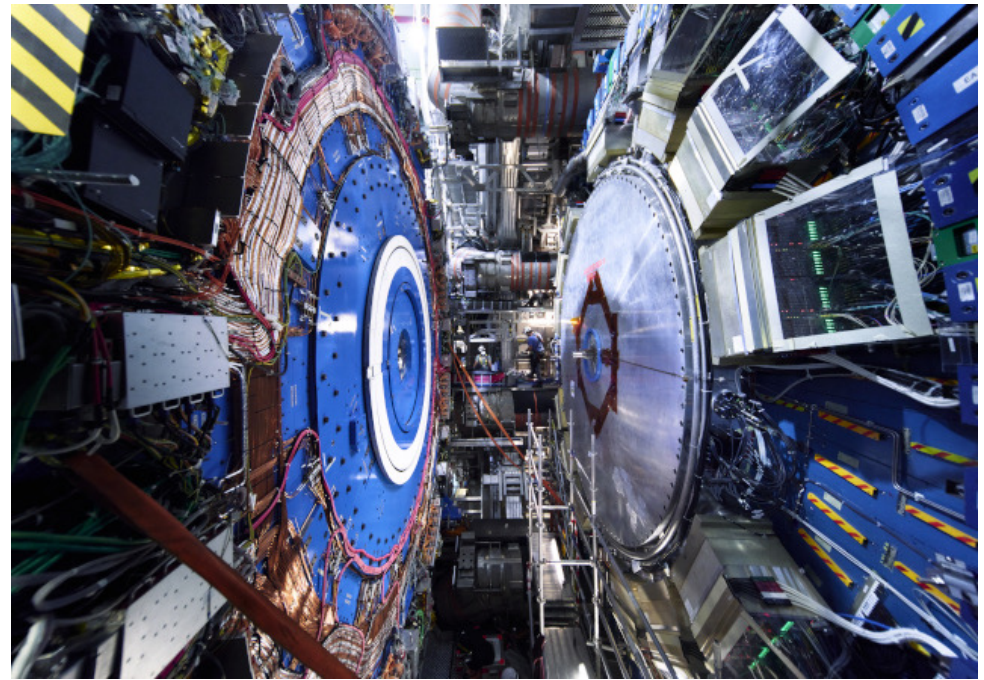


# Assembly and surface commissioning





# Installation in ATLAS





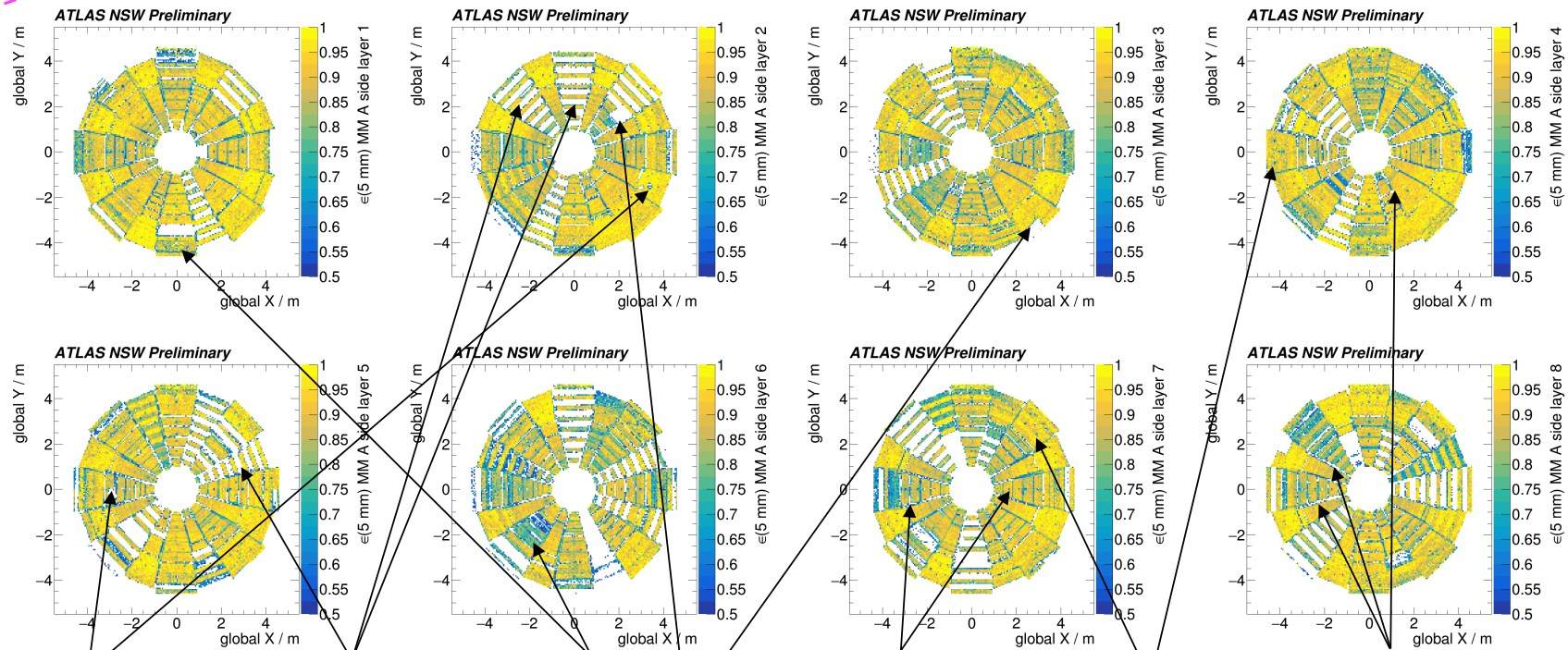
# Status of NSW

## NSW Preliminary Performance: MM Single layer efficiency



Inclusive of all inefficiencies (detector, HW, DAQ)

±5 mm

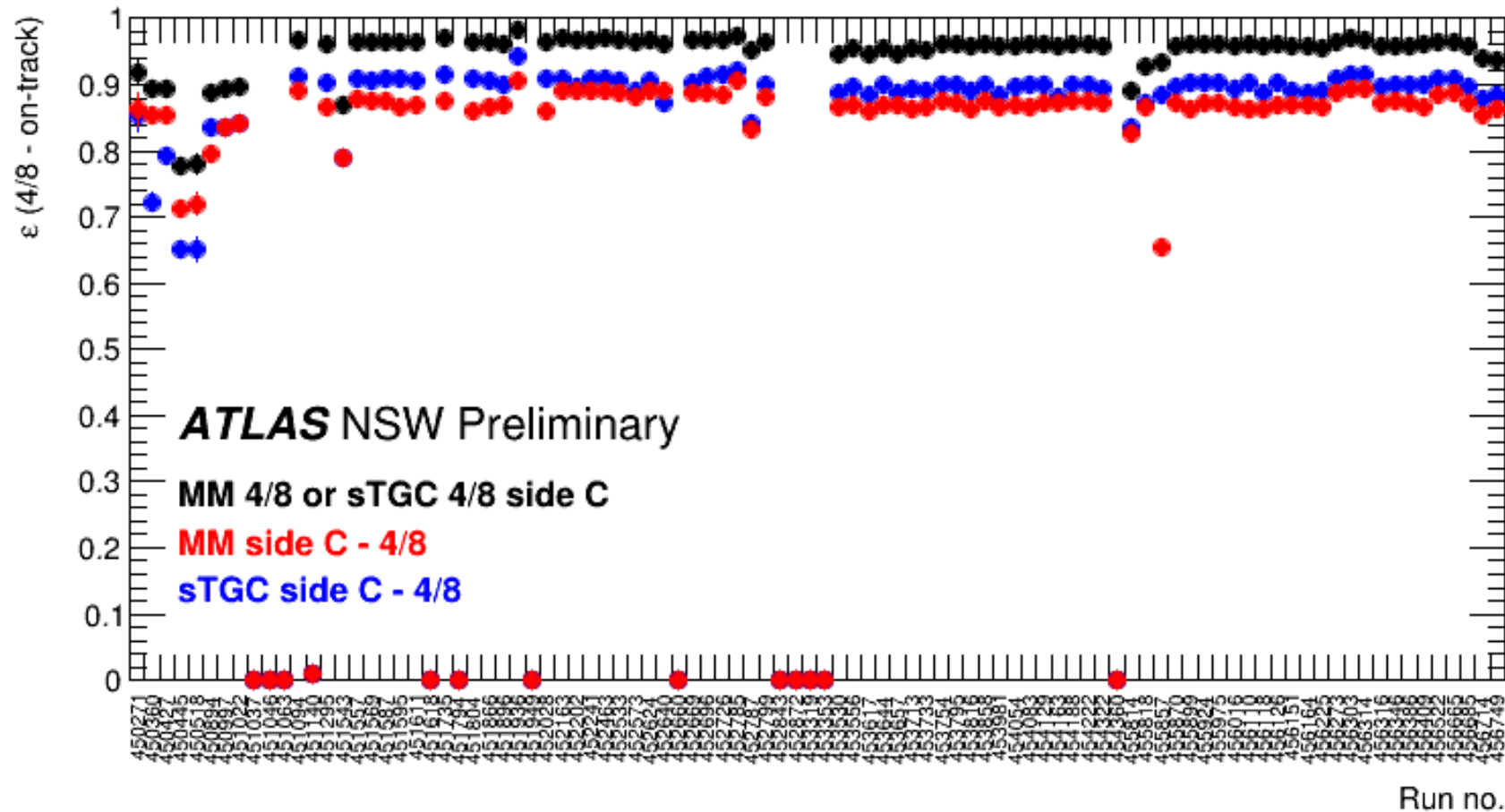


- Local defects
- Readout board problems or LV board problems (from these maps is impossible to distinguish between the problems)
- Lower HV
- Half HV channel disconnected
- Longitudinal passivations (non active area)
- Local passivation (interconnections screws, non active area)
- Side passivations (non active area)

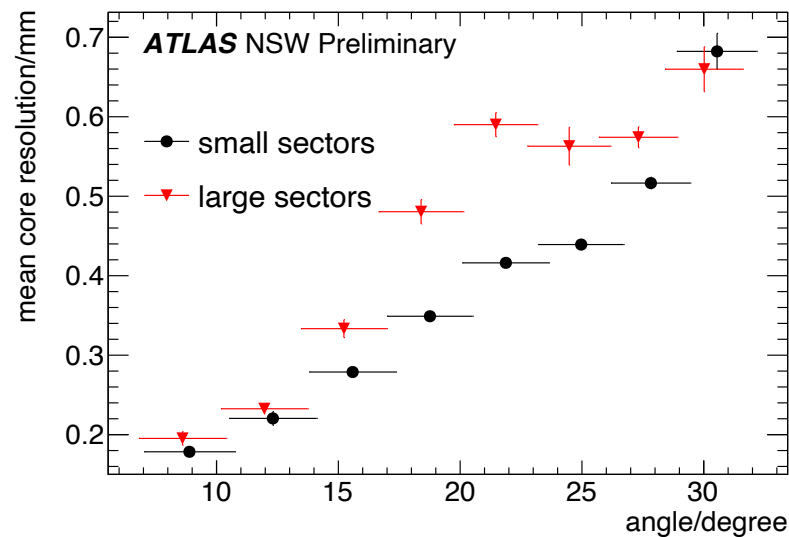


# Status of NSW

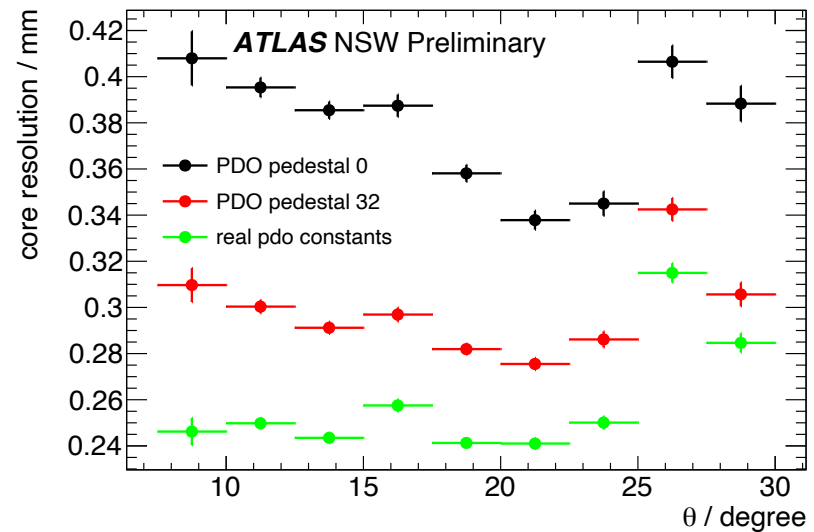
## Segment efficiency



Redundancy helped



(a)

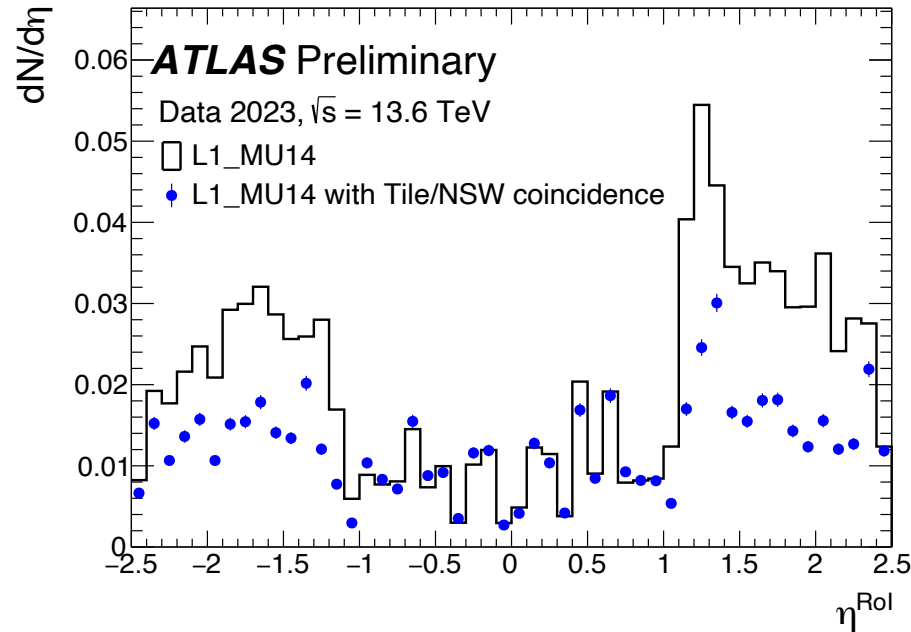


(b)

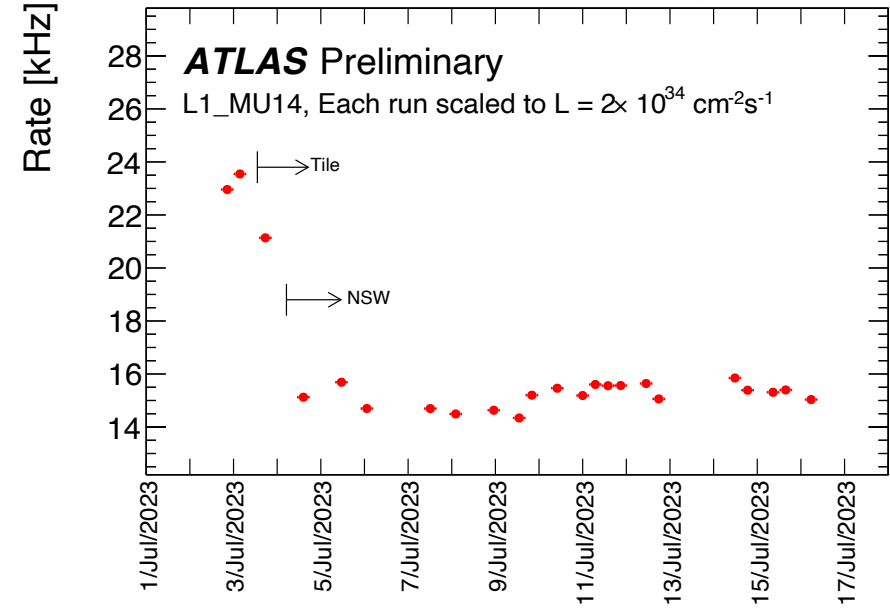
Figure 3: (a) Micromegas position resolution extracted by comparing the cluster position on two neighbouring layers corrected by the track angle for small and large sectors [7]. (b) sTGC resolution for different values of the charge calibration of the readout electronics. The resolution is calculated between the reconstructed cluster position on each layer and the muon track reconstructed with the full ATLAS detector [7].

# Status of NSW

Reduction of L1 rate using 100/144 sTGC pad-tower.  
sTGC strips and MM trigger will follow.



(a)



(b)

Figure 4: (a) The pseudorapidity ( $\eta$ ) distribution of the level-1 (L1) Region-of-Interests, which fulfill the primary L1 muon trigger with a threshold of the transverse momentum of 14 GeV before and after the deployment of the Tile and NSW coincidences in the L1 trigger decisions in 2023 data. Only the sTGC-Pad readout is used for the NSW coincidence of the track candidates [8]. (b) The trigger rate of the primary L1 muon trigger with a threshold of the transverse momentum of 14 GeV (L1\_MU14), scaled to the instantaneous luminosity of  $2 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ , as a function of time in 2023. The rate reduction after the reactivation of the Tile coincidence and after the inclusion of the NSW Pad Trigger was measured to be  $\sim 2 \text{ kHz}$  and  $\sim 6 \text{ kHz}$  respectively [8].

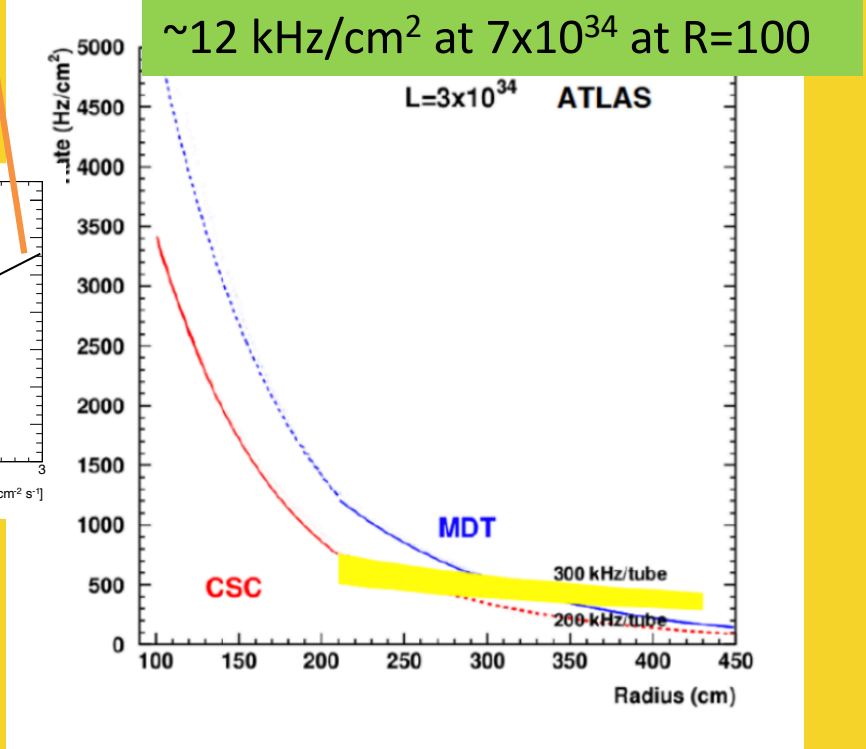
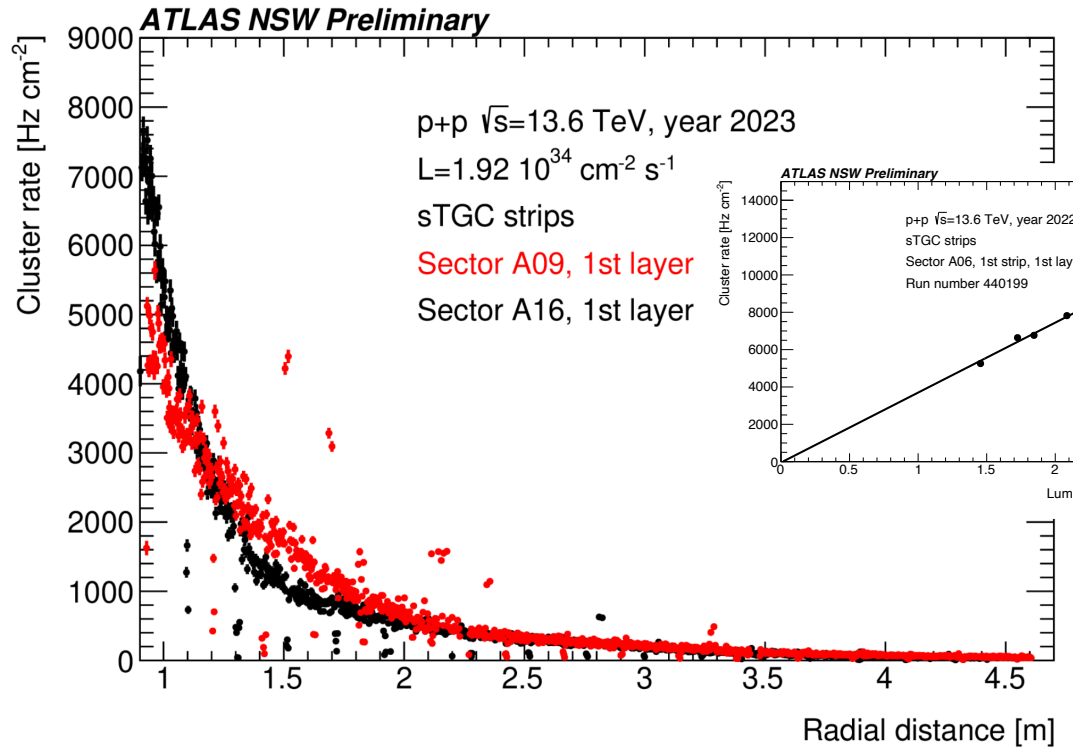
# Covern background hit rate

~25 kHz/cm<sup>2</sup> at 7x10<sup>34</sup> at R=100



2023 data

NSW TDR ATLAS



sTGC rate values from p+p data at R=100 cm:  
 small sector: ca. 7000 Hz cm<sup>-2</sup>  
 large sector: ca. 700 Hz cm<sup>-2</sup>

The rate at R=100 m, from the CSC based estimate of  
 from  $L=3 \cdot 10^{34}$   
 Hz cm<sup>-2</sup>

Measured hit rate is higher than initially considered, but it is still OK.

Different detector sensitivity, thinner radiation shielding, ...

sTGC rate  
 small sector: ca. 700 Hz cm<sup>-2</sup>  
 large sector: ca. 750 Hz cm<sup>-2</sup>

measurements  
 extrapolated to 14 TeV and to  $L=1.92 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$  is ca.  
 750 Hz cm<sup>-2</sup>

