

Semi-Annihilating Wino-Like Dark Matter

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THE UNIVERSITY OF
MELBOURNE



CoEPP
ARC Centre of Excellence for
Particle Physics at the Terascale

1507.xxxxx w/ Andrew Spray

About Me

- B.S. in Physics and Mathematics
Peking University, 2004
- Ph.D. in Physics
University of California at Davis, 2010
Advisor: Markus A. Luty
- 1st Postdoc
Shanghai Jiao Tong University, 2010-2012
- 2nd Postdoc
The University of Melbourne, 2012-present

Table of Contents

Dark Matter Symmetries and Semi-Annihilation

Semi-Annihilating Fermions

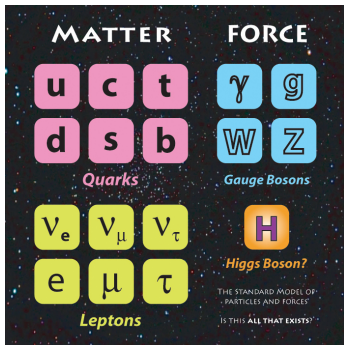
Relic Density

Other Constraints

Summary

The Two Standard Models

- Particle Physics

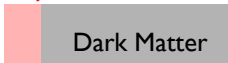


Higgs found in 2012

Ordinary $\Uparrow$ Matter



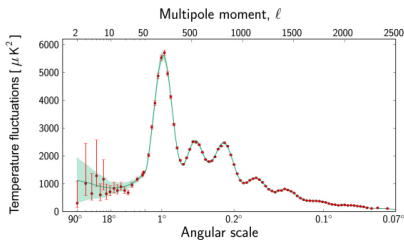
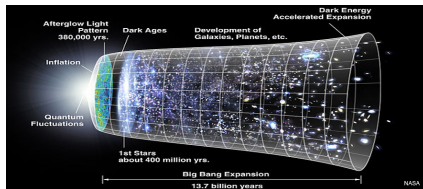
Dark Matter



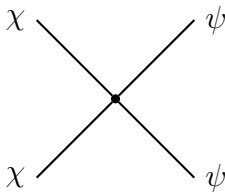
Dark Energy



- Cosmology



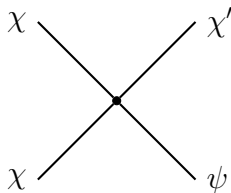
- Traditional: stabilize dark matter with Z_2 symmetry
 - Supersymmetry: R-Parity
 - Little Higgs: T-Parity
 - Extra dimensions: KK parity
 - Inert doublet model
- DM candidate: the lightest of Z_2 odd, single component
- The generic picture



Dark Matter Symmetries

Beyond- Z_2

- Almost any other symmetries
- Semi-Annihilation (D'Eramo and Thaler, 1003.5912)
- Multi-component dark matter
- More generic picture



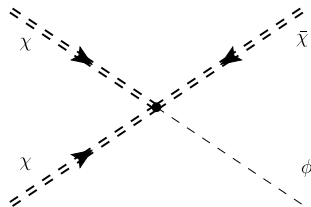
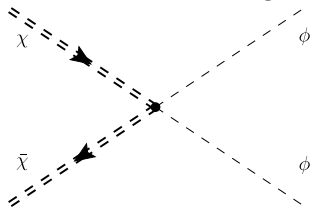
Semi-Annihilation

1003.5912

- A Z_3 toy model with pure scalar DM

$$\mathcal{L}_{Z_3} = m_\chi^2 \chi^\dagger \chi + a_1 \chi^\dagger \chi \phi + a_2 \chi^\dagger \chi \phi^2 + a_3 \chi \chi \chi \phi - V(\phi)$$

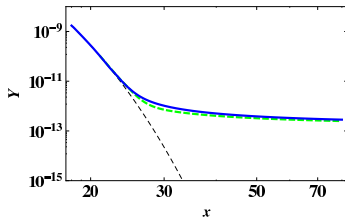
- χ : complex scalar with Z^3 charge 1
- ϕ : real scalar with Z^3 charge 0
- Two annihilation diagrams:



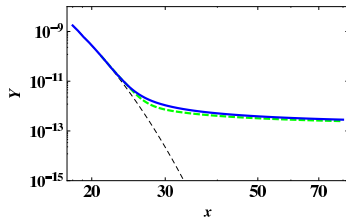
Semi-Annihilation

The relic density

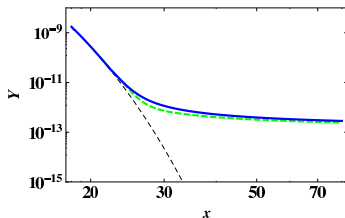
$$\eta_a=0.61, \eta_s=0, \Omega h^2 \simeq 0.1$$



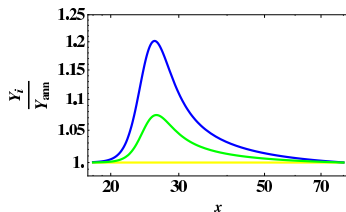
$$\eta_a=\eta_s=0.47, \Omega h^2 \simeq 0.1$$



$$\eta_a=0, \eta_s=0.72, \Omega h^2 \simeq 0.1$$

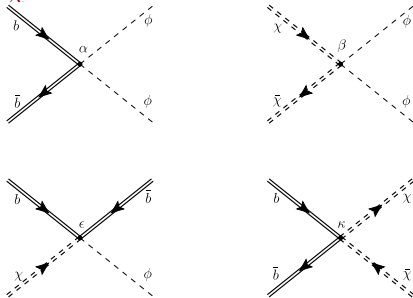


$$Y_i/Y_{\text{ann}} \text{ vs } x$$



Multi-component model

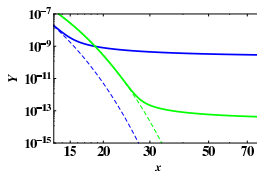
$bb\chi$ model



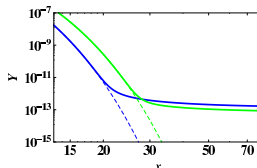
- $b, \bar{b}$ vector-like, $U(1)$, ± 1
- χ, ϕ , complex/real scalar, $-2/0$
- α, β , self-annihilation
- ϵ , semi-annihilation
- κ , dark matter conversion
- example, $\alpha = 0.03, \beta = 1.5$

$$m_\chi = 0.8 \text{ TeV}, m_b = 1 \text{ TeV}$$

$$\kappa = 0, \epsilon = 0, \Omega h^2 \simeq 100$$



$$\kappa = 2, \epsilon = 0, \Omega h^2 \simeq 0.1$$



$$\kappa = 0, \epsilon = 1.5, \Omega h^2 \simeq 0.1$$

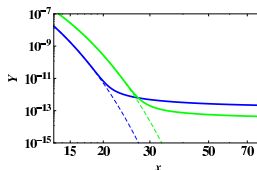


Table of Contents

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Semi-Annihilating Fermions

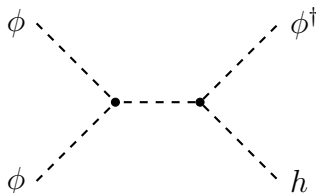
Relic Density

Other Constraints

Summary

Previous Models

- Scalar(or vector) dark matter
 - Batell et al 1007.0045, 1103.3053; D'Eramo et al, 1210.7817;
 - Belanger et al 1112.4491, 1211.1014, 1403.4960; Ivanov and Keus, 1203.3426;
 - Exception: Aoki and Toma, 1405.5870.
- Reason: simplest symmetries Z_3 and Z_4 to have χ^3 and χ^4
 - Bosonic cubic and quartic terms are renormalizable.
 - Scalar semi-annihilating models are simpler.



With Fermions

A Minimal Model

- Minimal flavor-neutral case

- Z_4 symmetry, Dirac fermion Ψ and real scalar ϕ
 $\Psi \sim (1, 1)_0 \quad G_{Z_4}(\Psi) = 1 \quad \phi \sim (1, 1)_0 \quad G_{Z_4}(\phi) = 2$
- Ψ only couples to ϕ , not SM and thus completely invisible
- Phenomenology reduces to scalar singlet model.
Notable exception: galactic center excess

- Minimal **interesting** case

- Z_4 symmetry, Dirac fermion Ψ and real scalar ϕ
 $\Psi \sim (1, 3)_0 \quad G_{Z_4}(\Psi) = 1 \quad \phi \sim (1, 1)_0 \quad G_{Z_4}(\phi) = 2$
- Fermion: Dirac "wino"
- Scalar: Higgs-portal singlet
- Relevant couplings:

$$\mathcal{L} \sim \frac{1}{2} \lambda \phi^2 H^\dagger H + y(\phi \bar{\Psi}^c \Psi + h.c.)$$

Table of Contents

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Relic Density

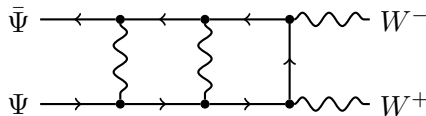
Other Constraints

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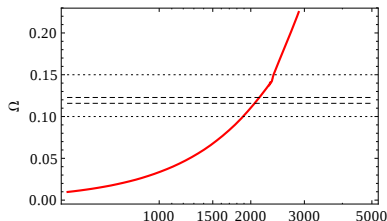
Fermion Relic Density

Sommerfeld Enhancement

- Fermion annihilates via gauge couplings
 - Sommerfeld effect: non-perturbative
Hisano et al, hep-ph/0412403, hep-ph/0610249
Cirelli et al, 0706.4071; Feng et al, 1005.4678



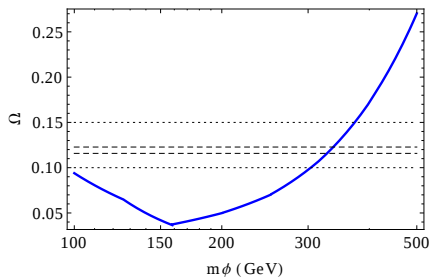
- Resonance at 2.4 TeV
- Relic density for $m_{\Psi} \sim 2$ TeV, lower than Wino



Scalar Relic Density

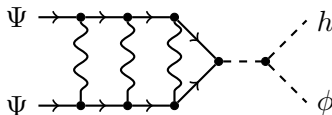
Cline et al, 1306.4710

- Dominant: s -channel Higgs portal into SM particles
- Sub-dominant: anni. into hh
- High mass: relic density for $m_\phi \sim (3 \text{ TeV}) \times \lambda$
- Low mass: more complex(WW^* , ZZ^* decays)

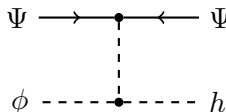


Semi-Annihilation

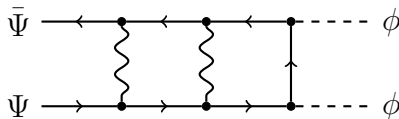
- Two semi-annihilation channels
 - $\Psi\Psi \rightarrow \phi h$: Sommerfeld enhanced



- $\Psi\phi \rightarrow \bar{\Psi}h$: only changes the scalar number



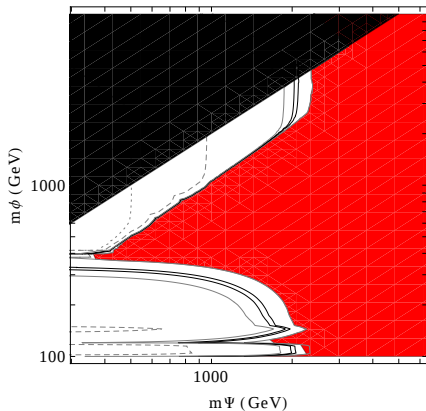
- DM exchange $\Psi\Psi \leftrightarrow \phi\phi$: Sommerfeld enhanced



Benchmark Scenarios

- $y = 0.1, \lambda = 0.1$
- Red: $\Omega h^2 > 0.15$
- Black: scalar unstable

$y = 0.1$



scalar and fermion density

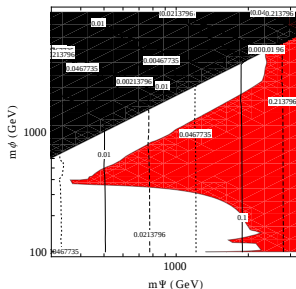
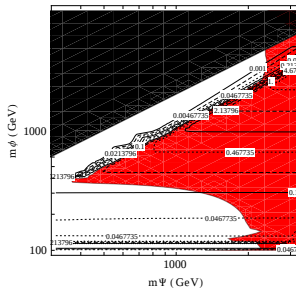


Table of Contents

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Semi-Annihilating Fermions

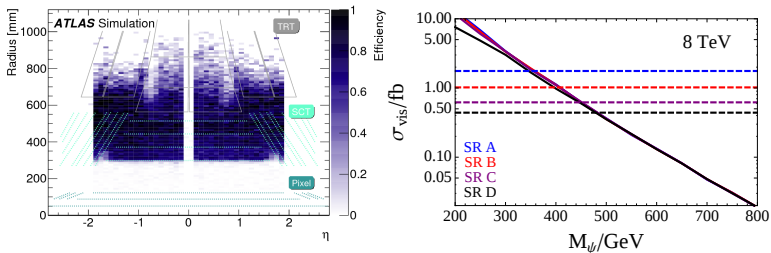
Relic Density

Other Constraints

Summary

Collider Search

- Mono-jet puts constraints on m_ϕ , but very weak
- Disappearing tracks $pp \rightarrow W \rightarrow \Psi^\pm \Psi^0$, $\Psi^\pm \rightarrow \pi^\pm \Psi^0$
 - $p_T > 90$ GeV, $E_T^{miss} > 90$ GeV, $\Delta\phi_{min}^{jet-E_T^{miss}} > 1.5$
 - lepton veto
 - disappearing-track: $\Psi^\pm$ decays after SCT and before TRT
 - four signal regions with $p_T > 75, 100, 150, 200$ GeV

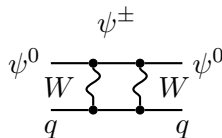
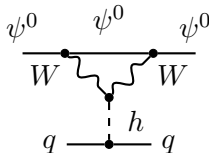
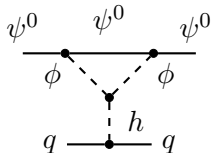


- $m_\Psi > 480$ GeV

Direct Detection and Indirect Detection

- Direct detection

- Fermion: loop suppressed



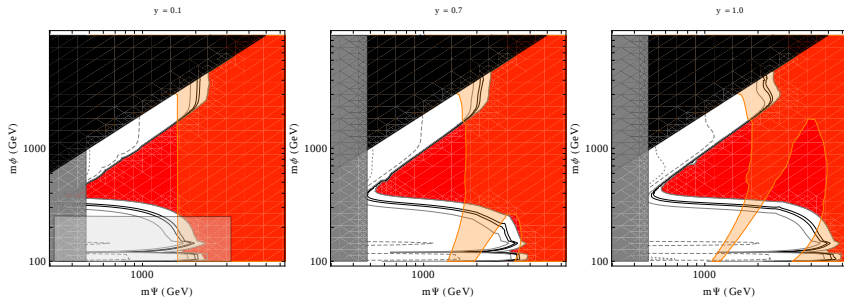
- Scalar: Higgs portal

$$m_\phi > (300 \text{ GeV}) \times (\lambda/0.1)^2 \times (\Omega_\phi/0.112)$$

- Indirect detection

- HESS: Gamma ray lines from fermion annihilation at the galactic center
- Sommerfeld enhanced, assume NFW profile

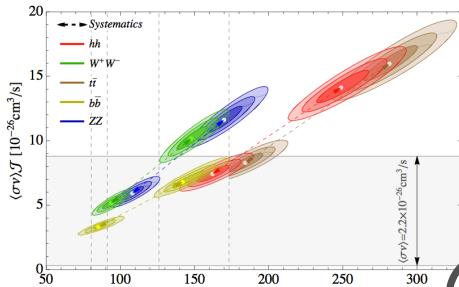
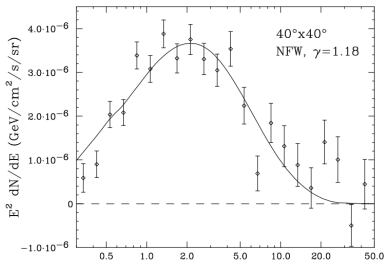
Constraint Plots



Gray: LHC Orange: HESS Lightgrey: Direct detection
Upper and lower bounds
Semi-annihilation and multi-component region remains.

Galactic Center Excess

- Broad excess in Fermi γ ray data
 - Hooper et al, 1010.2752, 1110.0006, 1305.0830;
 - Abazajian and Kaplinghat, 1207.6047;
 - Boehm et al, 1401.6458; Daylan et al, 1402.6703.
- Fermi collaboration: only a talk
- Alternative explanation: pulsars
Mirabal, 1309.3428; Yuan and Zhang, 1404.2318; ...
- Perhaps dark matter annihilation
Agarwal et al, 1411.2592; Colore et al, 1411.4647



Fit

- The simplest approach: pure scalar annihilation
- Assumption: $\Omega_\phi \approx \Omega_{\text{DM}}$
- Fit spectrum for $\phi\phi \rightarrow \text{SM}$
- Need semi-annihilation to get right relic density
- m_Ψ, y

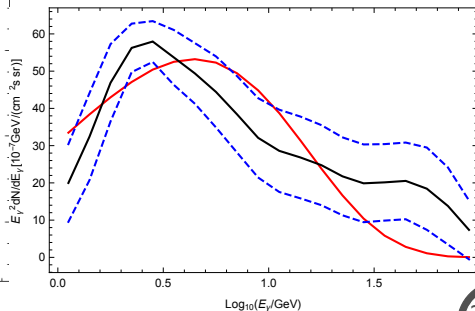
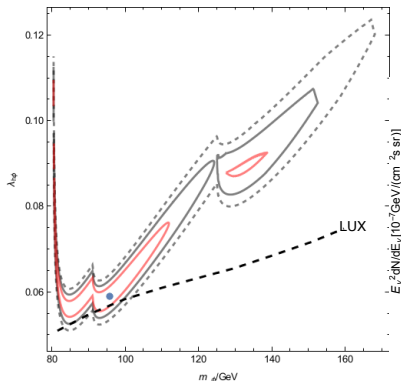


Table of Contents

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- Generic feature of dark sectors: semi-annihilation
- Multi-component: necessary for fermion models
- Simple models: wino-like Ψ and Higgs portal
- Limits from relic density and other observations. It can be excluded
- Fermi fit need semi-annihilation