

Total charm cross-section in Au+Au 200GeV from STAR

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Charm hadron measurement at Au+Au 200GeV

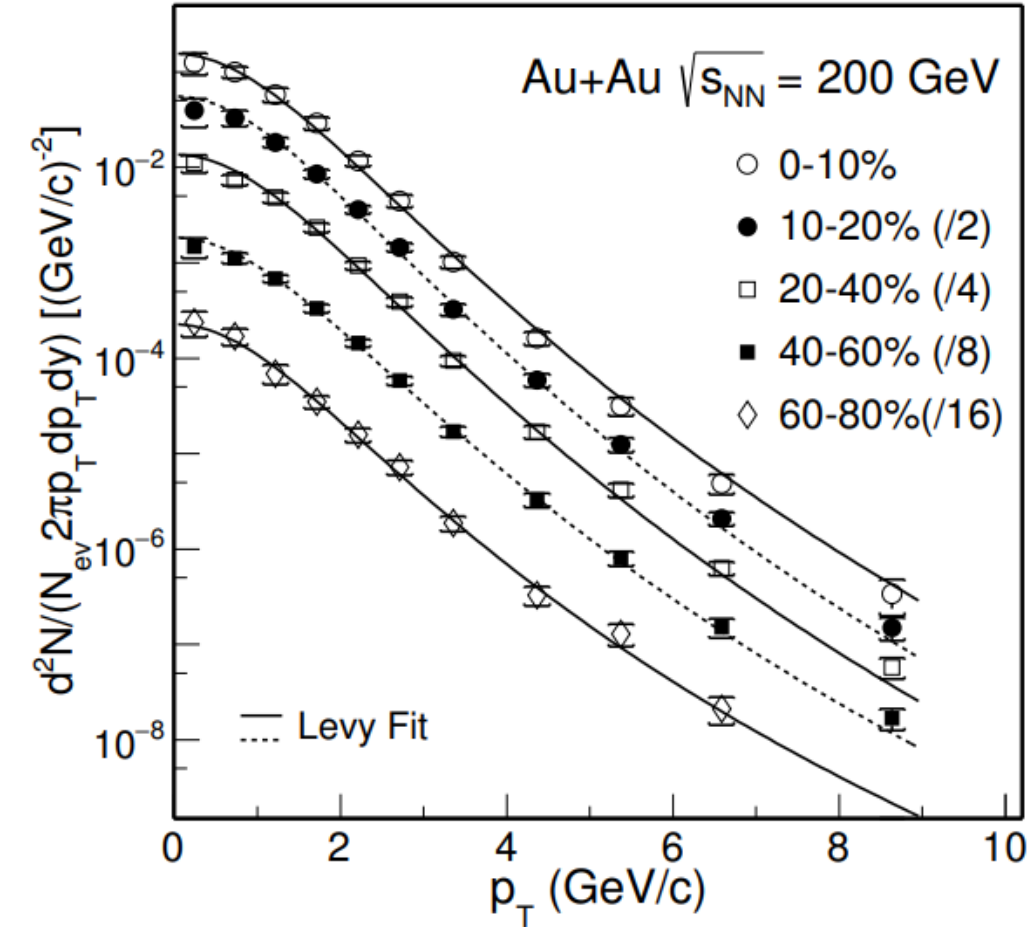
Ground state charm hadrons production(with HFT) at Au+Au 200GeV from STAR:

- D0: run14, published PhysRevC.99.034908 (2019)
- Lc: run14&run16, published Phys. Rev. Lett. 124, (2020) 172301
- Ds:run14&run16, published Phys. Rev. Lett. 127, (2021) 092301
- D $\pm$:run14&run16,
https://www.star.bnl.gov/protected/heavy/vanekjan/Dpm_web/Home.html
- All primary Ground state charm hadrons are already published or in publish step.
So we want to carry out a total charm production cross-section results in D $\pm$ paper

Previous calculation by Xiaolong Chen

https://drupal.star.bnl.gov/STAR/system/files/charmCrossSection_190627.pdf

D⁰ production cross-section:



Published D⁰ data cover whole p_T range (0 GeV/c to 10 GeV/c)

$$\frac{d\sigma}{dy} \Big|_{y=0} = \int \frac{d^2N}{dy dp_T} \Big|_{y=0} dp_T * \frac{\sigma_{pp}}{N_{collision}}$$

$\int \frac{d^2N}{dy dp_T} \Big|_{y=0} dp_T$ is calculated by sum all measured dN/dy in each p_T

$$\text{bin: } \int \frac{d^2N}{dy dp_T} \Big|_{y=0} dp_T = \sum_{p_T \text{ bin}} \frac{dN_{measured}}{dy} \Big|_{y=0}$$

$$\sigma_{pp} = 42 \text{ mb}$$

$N_{collision}$ taking from run14&16 centrality calculation

(the σ_{pp} and $N_{collision}$ value will give an 8% uncertainty on all charm hadron cross-section calculation)

<https://www.star.bnl.gov/protected/heavy/xgn1992/Centrality/Run2014/>

D⁰ production cross-section:

- For the uncertainty calculation, assuming the stat. err. in each pT bin are uncorrelated and sys. err. are fully correlated. The results shows in the table below:

Centrality:	$\frac{d\sigma}{dy} _{y=0} (\mu b)$	Stat. Err.	Sys. Err.
0-10%	38.1894	1.1948	2.1799
10-40%	39.0188	0.5715	1.1132
40-80%	36.6356	0.6765	1.4079
0-80%	39.4365	0.4584	5.3178

$0.0 < p_T < 8.0 \text{ GeV}/c$

- Considering both statics and consistence with other charm hadron measurement, we choose 10-40% centrality results **$39.0 \pm 0.6 \text{ (stat.)} \pm 1.1 \text{ (sys.)}$** to calculate the total charm cross-section

D⁰ production cross-section: Fit to different model

- Using different model functions to describe D⁰ p_T spectra in low p_T range
- Function choosing:

Levy :

$$\frac{d^2 N}{2\pi p_T dp_T dy} = \frac{1}{2\pi} \frac{dN}{dy} \frac{(n-1)(n-2)}{nT(nT + m_0(n-2))} \left(1 + \frac{\sqrt{p_T^2 + m_0^2} - m_0}{nT}\right)^{-n}$$

Power-law:

$$\frac{d^2 N}{2\pi p_T dp_T dy} = \frac{dN}{dy} \frac{4(n-1)(n-2)}{2\pi(n-3)^2 \langle p_T \rangle^2} \left(1 + \frac{2p_T}{\langle p_T \rangle (n-3)}\right)^{-n}$$

Blast-wave:

$$\frac{dN}{p_T dp_T} = \frac{dN}{m_T dm_T} \propto \int_0^R r dr m_T I_0\left(\frac{p_T \sinh \rho}{T_{\text{kin}}}\right) K_1\left(\frac{m_T \cosh \rho}{T_{\text{kin}}}\right)$$

Tsallis:

$$E \frac{d^3 N}{dp^3} = gV \frac{m_T}{(2\pi)^3} \left[1 + (q-1) \frac{m_T}{T_1}\right]^{-q/(q-1)}$$

Boltzmann:

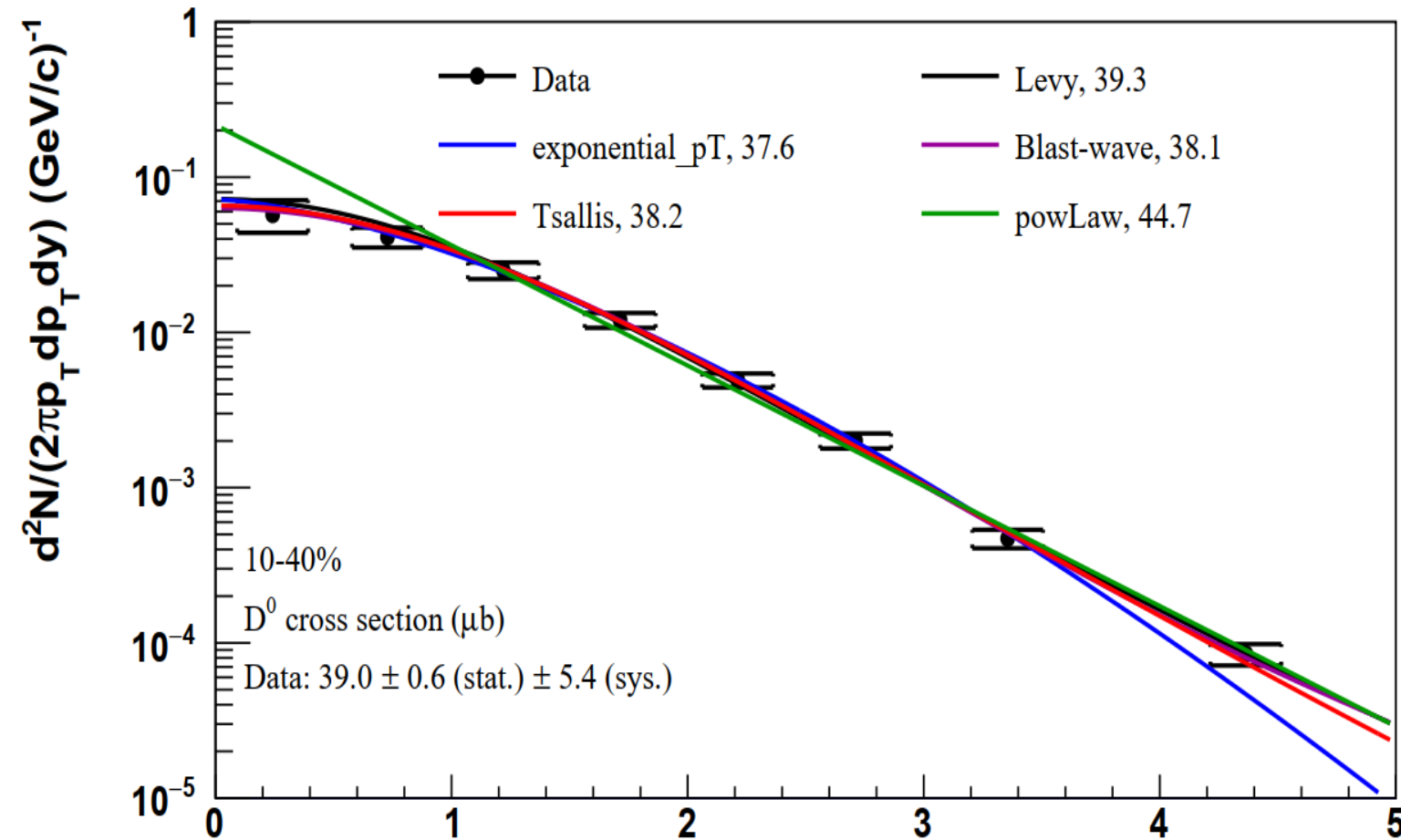
$$\frac{d^2 N}{dp_T dy} \propto m_T \exp(-m_T/T)$$

exponential (p_T, p_T-Gaus, p_T-trip):

$$\frac{dN}{p_T dp_T} = c_{p_T^{1.5}} \exp\left(-\frac{p_T^{1.5}}{T_{p_T^{1.5}}}\right), \quad \frac{dN}{p_T dp_T} = c_{p_T^2} \exp\left(-\frac{p_T^2}{T_{p_T^2}^2}\right), \quad \frac{dN}{p_T dp_T} = c_{p_T^3} \exp\left(-\frac{p_T^3}{T_{p_T^3}^3}\right)$$

D^0 cross-section: Fit to different model

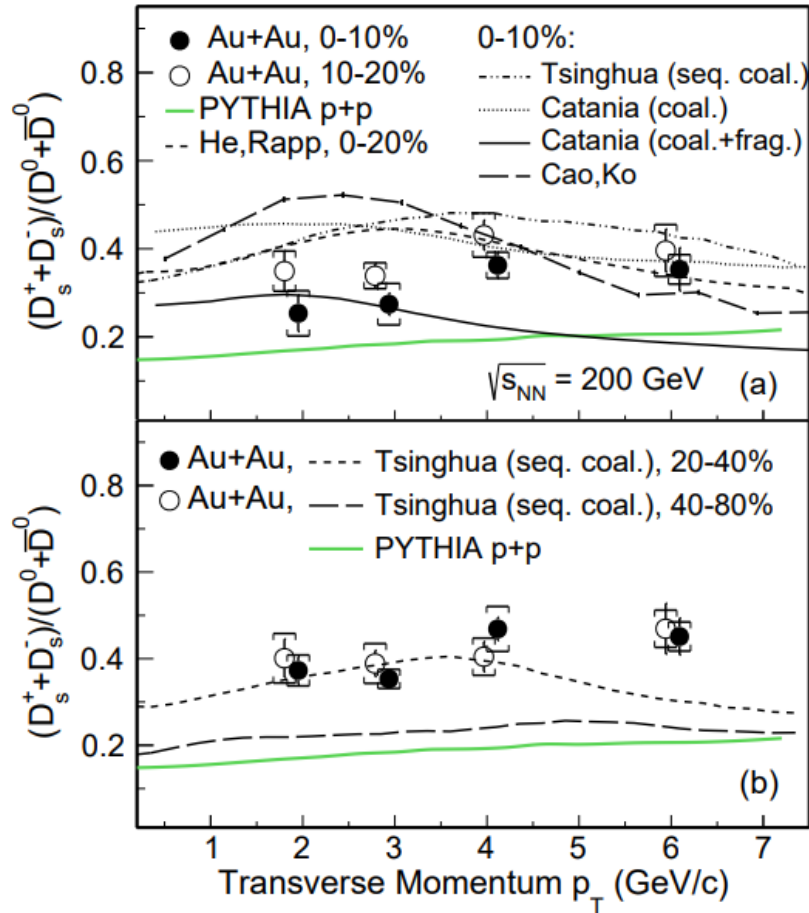
Number after function name means Integral cross-section in $p_T(0.0-10.0 \text{ GeV}/c)$



Boltzmann, Exponential p_T -Gaus, Exponential p_T -triple are failed to describe data.

Levy , Blast-wave, Tsallis and Exponential p_T and Power-Law functions can fit to the low p_T spectra, but Power-Law function overestimate 0-1 GeV yields.

D_s cross-section: extrapolate from D_s/D^0 ratio



D_s measurement for 10-40% cover p_T 1.0 to 8.0 GeV/c,

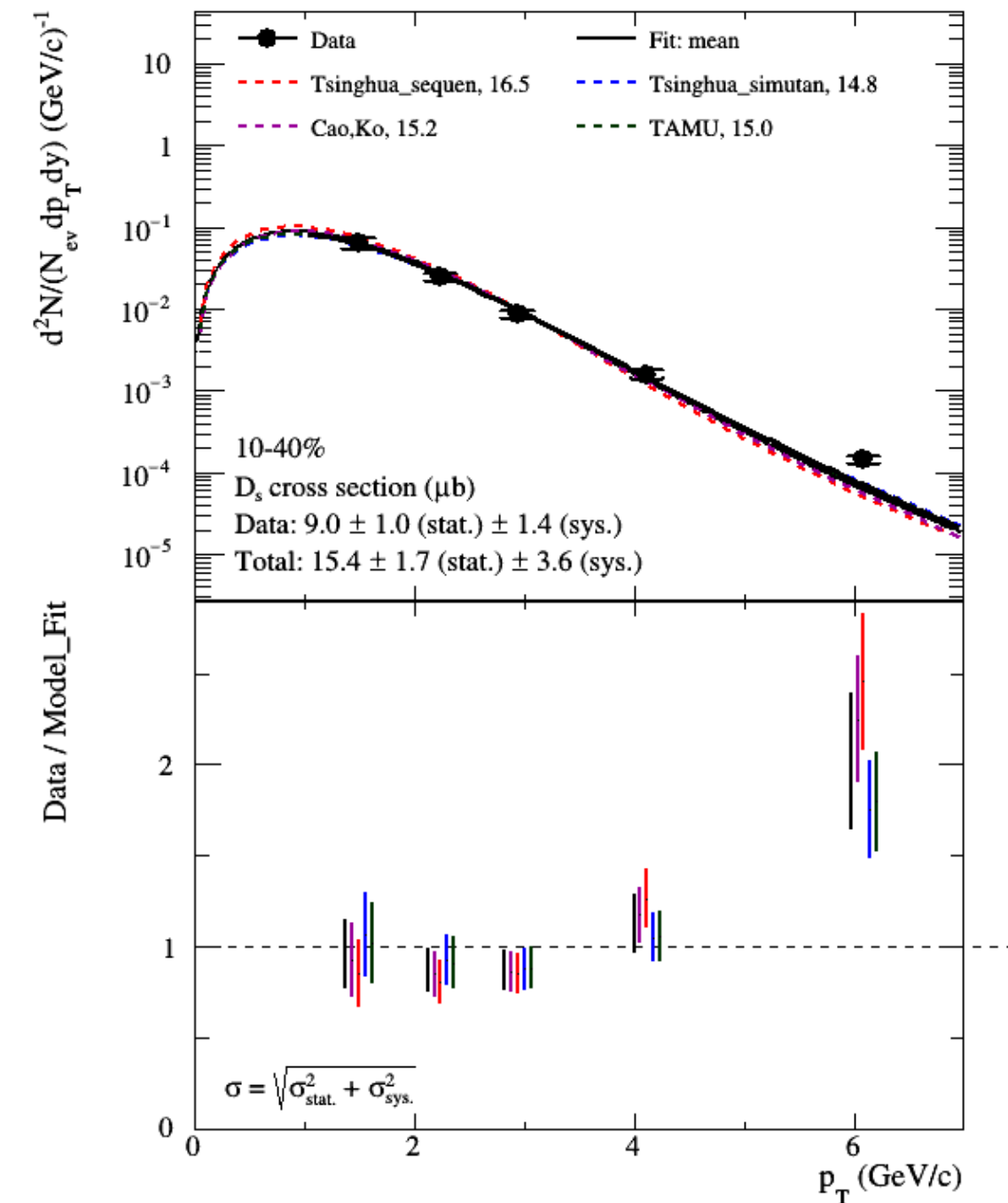
Using the same method as L_c to do the extrapolation down to 0 p_T :

$$\frac{dN(D_s)}{dp_T dy}(p_T) = \frac{dN(D^0)}{dp_T dy}(p_T) \times C_1 \frac{D_s}{D^0}(p_T)$$

$\frac{dN(D^0)}{dp_T dy}(p_T)$ is getting from levy fitting of D^0 10-40% centrality data (p_T 0-10 GeV/c).

$\frac{D_s}{D^0}(p_T)$ are getting from different model calculation

C_1 is a free parameter which allow the overall D_s/D^0 ratio could change



All model curves are using 10-40% results, expect for TAMU, which only have 0-20% centrality.

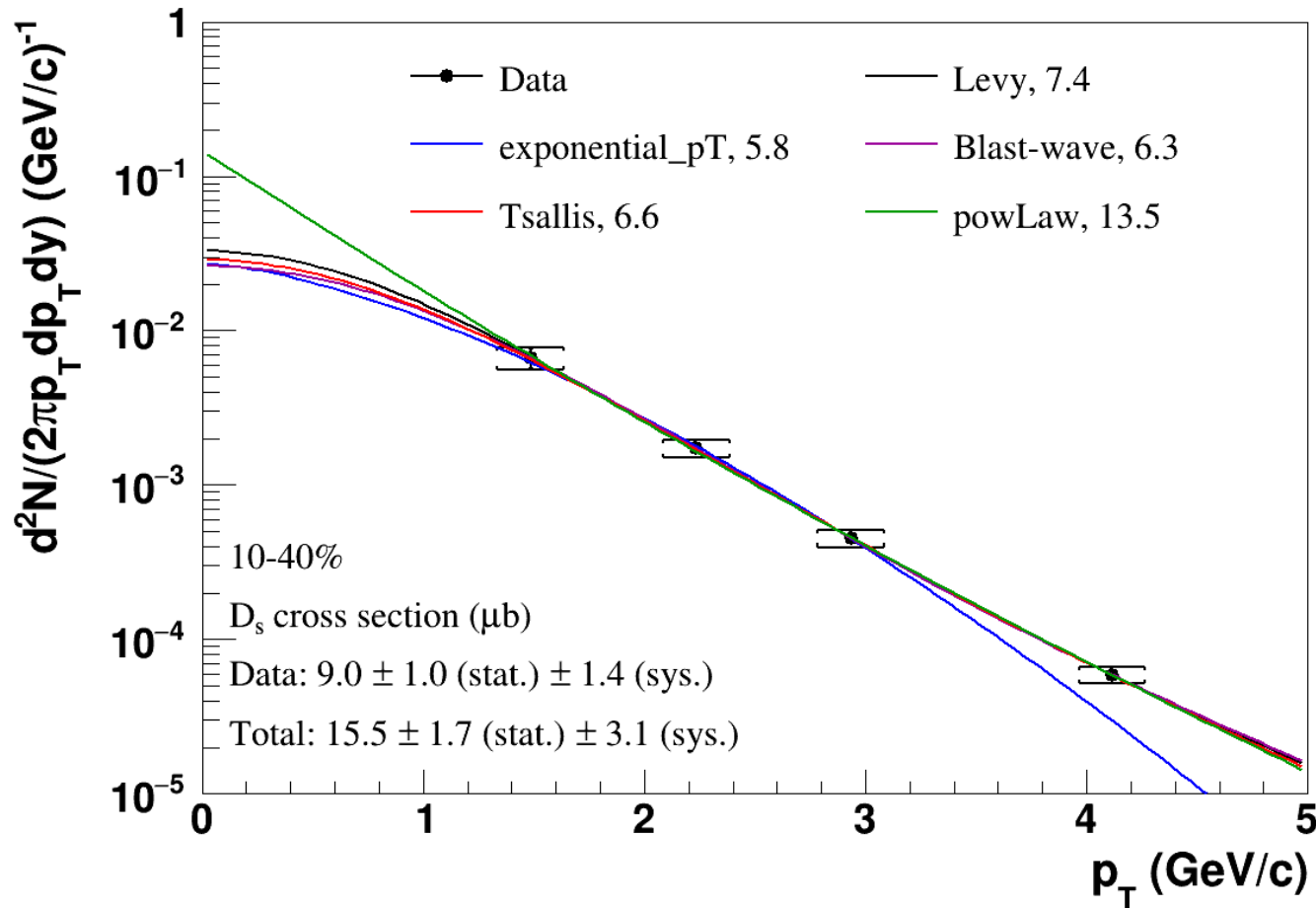
All D_s/D⁰ model are underestimated the 5.0-8.0 GeV data point

The final results are using the average between these four-model extrapolation (0-1 GeV) plus data (1-8.0 GeV). And the differences between models are quote as systemic errors.

Ds (10-40%) Total cross-section using this method is:
15.4 ± 1.7 (stat.) ± 3.6(sys.) 0.0 < p_T < 8.0 GeV/c

D_s cross-section: extrapolate from p_T spectra

Number after function name means Integral cross-section in $p_T(0.0-1.0 \text{ GeV}/c)$



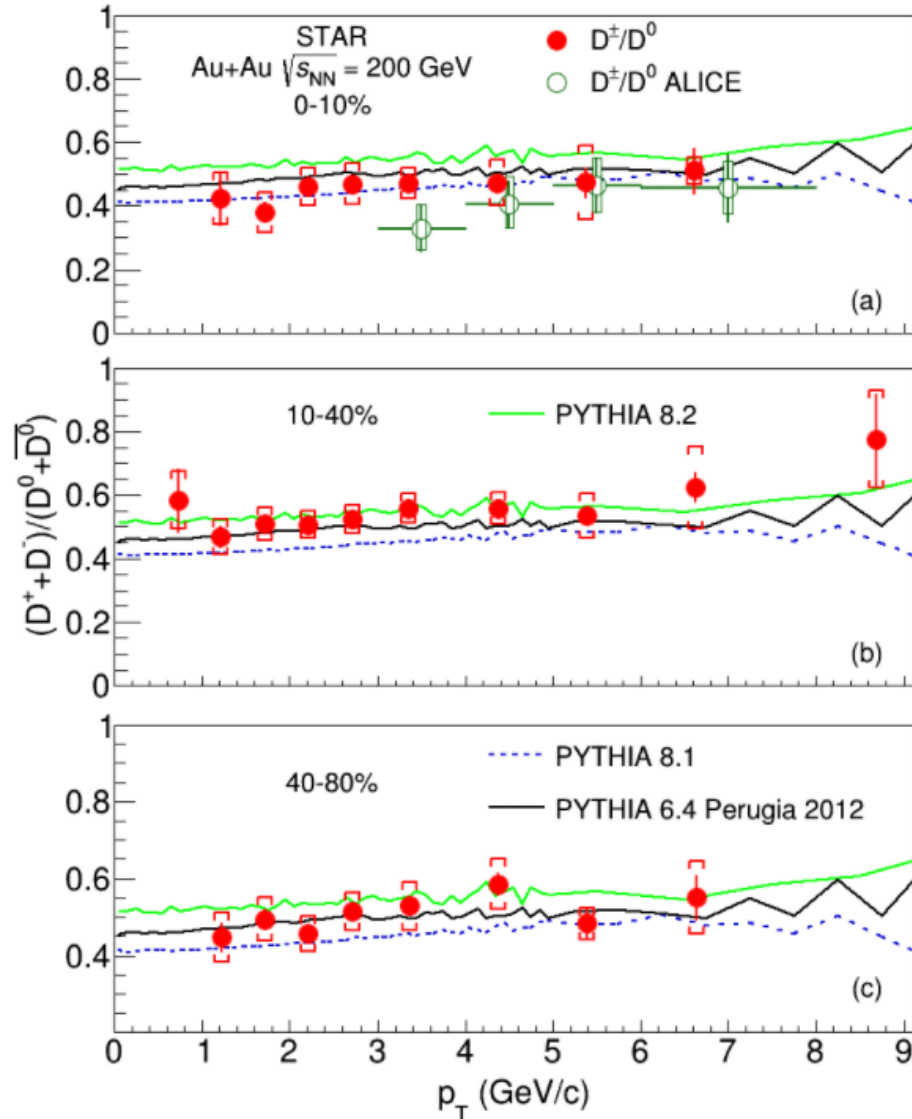
Boltzmann, Exponential p_T -Gaus, Exponential p_T -triple are failed to describe data.
Power-Law overestimate low p_T cross-section.

The final results are using the average between Levy, Blast-wave, Tsallis and Exponential p_T . And the differences are quote as systemic errors.

D_s (10-40%) Total cross-section using this method is: **15.5 ± 1.7 (stat.) ± 3.1 (sys.)**

The difference of the total cross-section between this and the previous method is very small

$D^\pm$ cross-section: extrapolate from $D^\pm/D^0$ ratio



$D^\pm$ measurement for 10-40% cover p_T 0.5 to 10.0 GeV/c,

Using the same method as L_c to do the extrapolation down to 0 p_T :

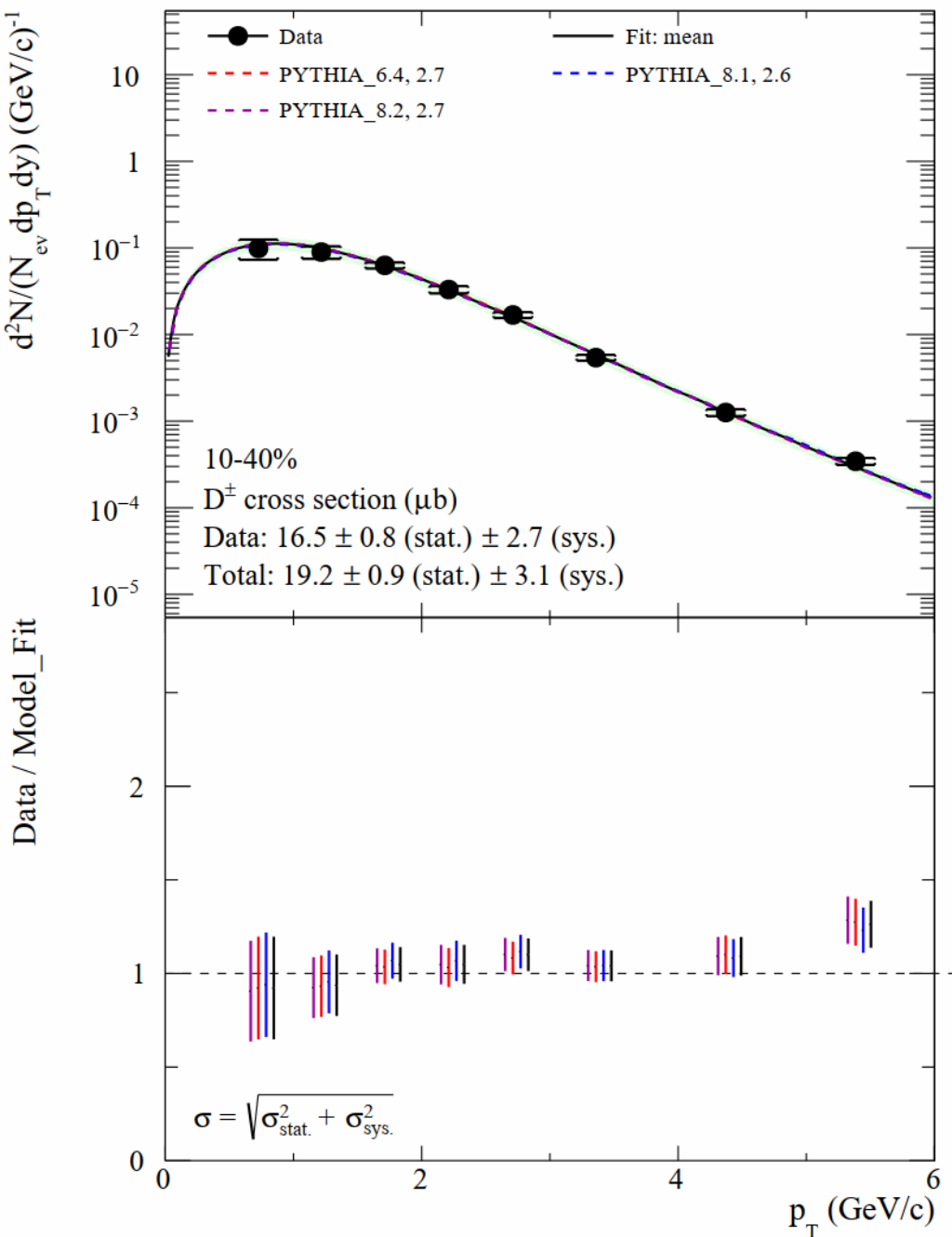
$$\frac{dN(D^\pm)}{dp_T dy}(p_T) = \frac{dN(D^0)}{dp_T dy}(p_T) \times C_1 \frac{D^\pm}{D^0}(p_T)$$

$\frac{dN(D^0)}{dp_T dy}(p_T)$ is getting from levy fitting of D^0 10-40% centrality data (p_T 0-10 GeV/c).

$\frac{D^\pm}{D^0}(p_T)$ are getting from different version of PYTHIA calculation

C_1 is a free parameter which allow the overall $D^\pm/D^0$ ratio could change

Number after function name means Integral cross-section in $p_T(0.0-0.5 \text{ GeV}/c)$



The final results are using the average between PYTHIA predictions and extrapolation to 0-0.5 GeV p_T range, then plus data 0.5 to 8.0 GeV p_T range. And the differences between models are quote as an additional systemic errors.

$D^\pm(10-40\%)$ Total cross-section using this method is:
 $19.2 \pm 0.9 \text{ (stat.)} \pm 3.1 \text{ (sys.)}$ $0.0 < p_T < 8.0 \text{ GeV}/c$

Λ_c production cross-section:

Λ_c measurement cover p_T 2.0 to 8.0 GeV/c

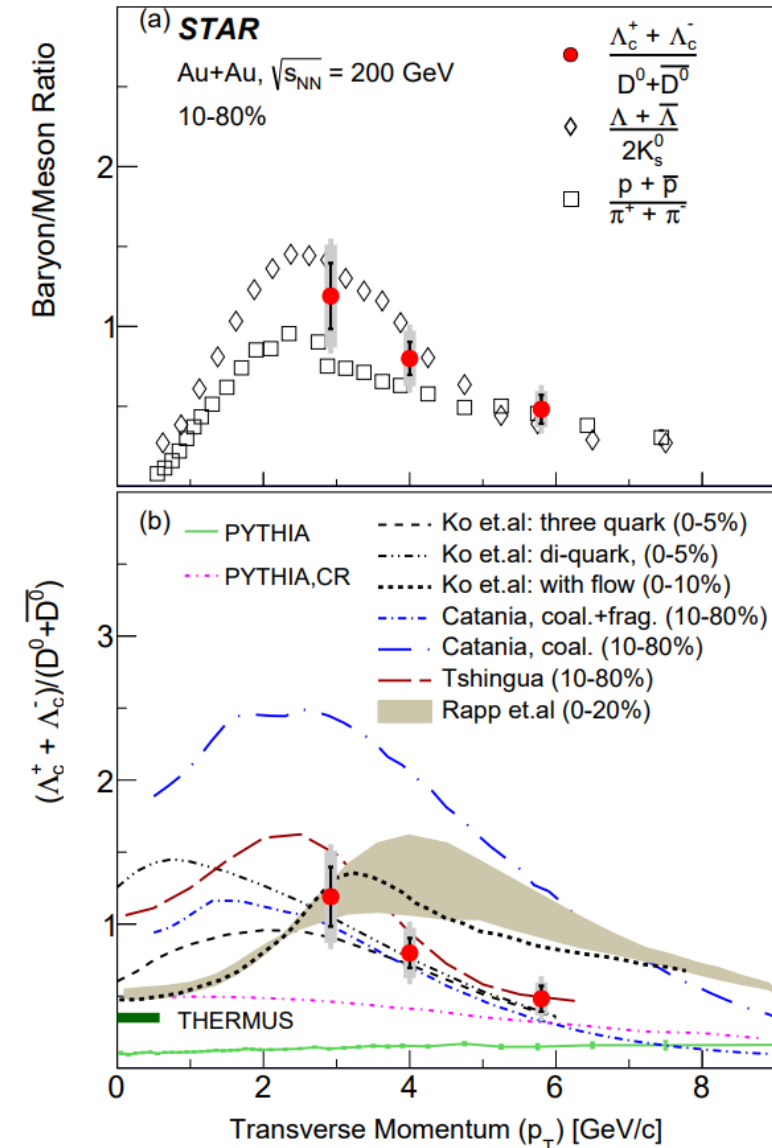
Using the measured Λ_c/D_0 ratio and D_0 spectra to do the extrapolation down to 0 p_T :

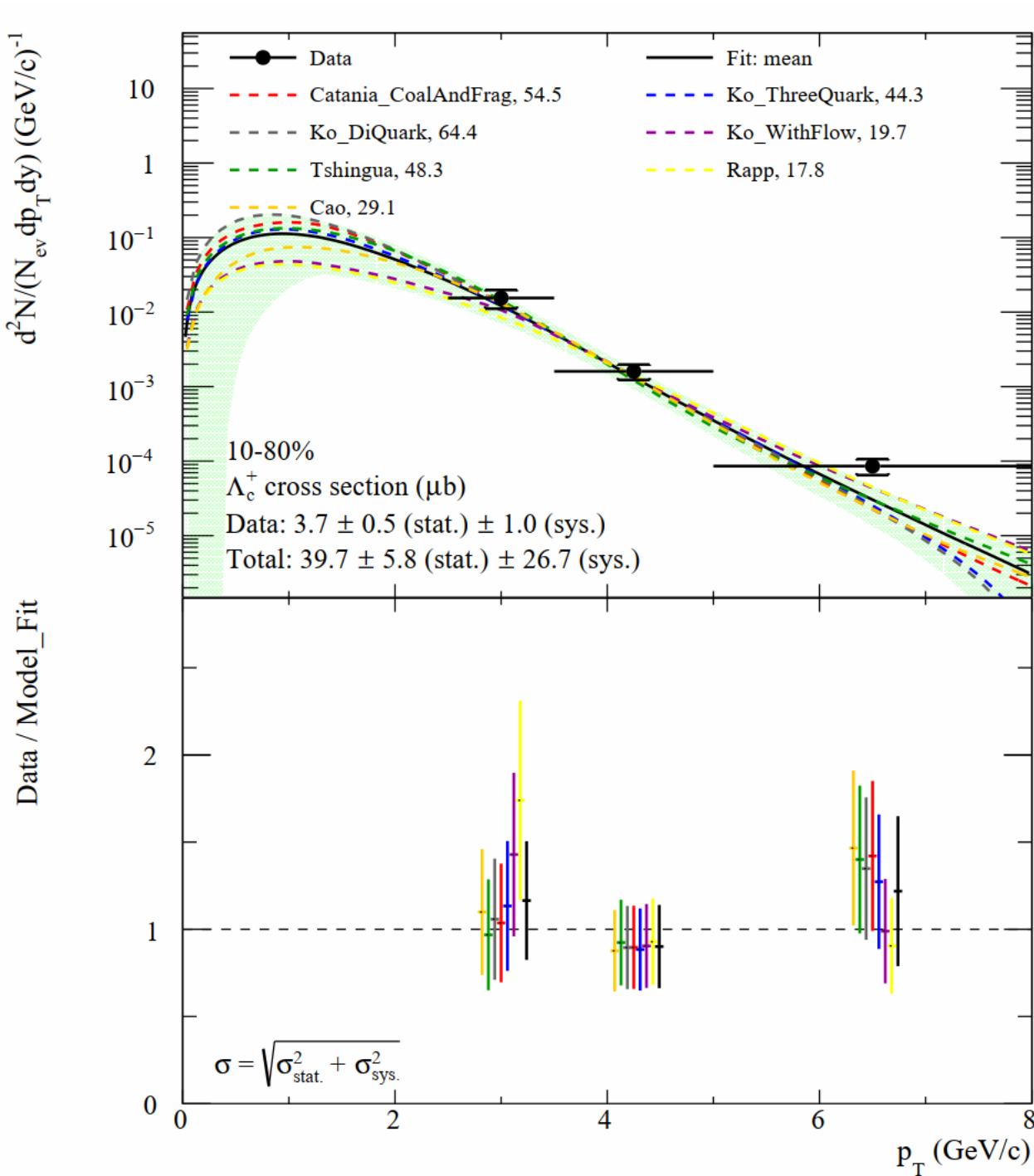
$$\frac{dN(\Lambda_c)}{dp_T dy}(p_T) = \frac{dN(D^0)}{dp_T dy}(p_T) \times C_1 \frac{\Lambda_c}{D^0}(p_T)$$

$\frac{dN(D^0)}{dp_T dy}(p_T)$ is getting from D^0 10-40% centrality data (p_T 0-10 GeV/c).

$\frac{\Lambda_c}{D^0}(p_T)$ are getting from different model prediction

C_1 is a free parameter which allow the overall Λ_c/D^0 ratio could change, and fitted to L_c spectra





Λ_c $d\sigma/dy$ in p_T 2.0 to 8.0 GeV/c are calculated by data and p_T 0.0 to 2.0 GeV/c are using the integrated average between these four-model. And the differences between models are quote as an additional systemic errors.

The bottom plot shows the ratio between Lc $d\sigma/dy$ data and model in 2.0 to 8.0 GeV/c

Lc (using 10-80% data) Total cross-section is:
 $39.7 \pm 5.8(\text{stat.}) \pm 26.7 (\text{sys.})$ $0.0 < p_T < 8.0 \text{ GeV/c}$

Total charm cross-section:

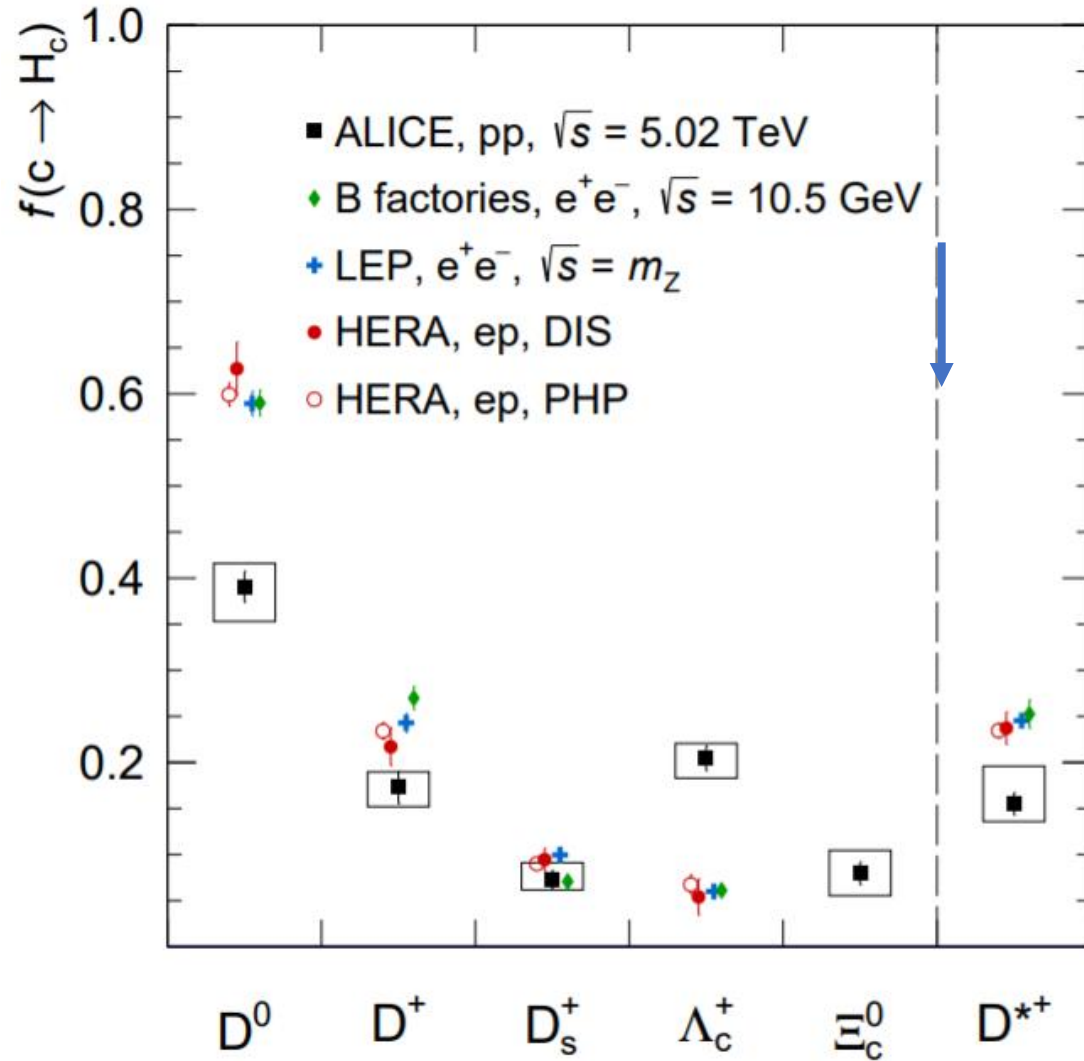
Collision System	Charm Hardon	Cross-section $\frac{d\sigma}{dy} _{y=0}$ (μb) (per nucleon-nucleon collision)
Au+Au 200 GeV (10-40% p _T 0-8 Gev/c)	D ⁰	39.0 ± 0.6 (stat.) ± 1.1 (sys.)
	D [±]	19.2 ± 0.9 (stat.) ± 3.1(sys.)
	D _s	15.4 ± 1.7 (stat.) ± 3.6(sys.)
	Λ _c ⁺ *	39.7 ± 5.8 (stat.) ± 26.7 (sys.)
	Total	113.3 ± 6.2 (stat.) ± 27.2 (sys.) *
P+P 200 GeV	Total	130 ± 30 (stat.) ± 26 (sys.)

* Λ_c⁺ results are using 10-80% centrality

* 8% uncertainty on σ_{pp} and N_{collision} are not included

Back up

Charm fragmentation at ALICE



Arxiv 2105.06335

Ground state charm hadron:

$D^0, D^+, D_s^+, \Lambda_c, \Xi_c^0, \Xi_c^+, \Omega_c, J/\psi$

$$f(c \rightarrow H_c) = \sigma(H_c) / \sigma_{\text{all}}$$

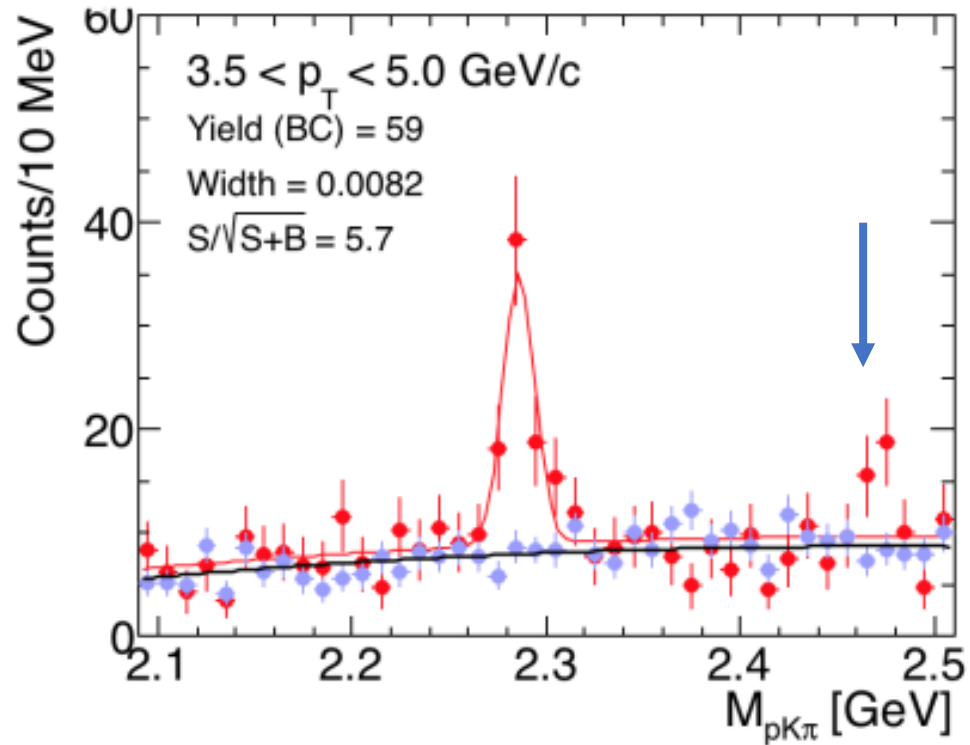
$$\text{HERA: } \sigma_{\text{all}} = \sigma(D^0) + \sigma(D^+) + \sigma(D_s) + \sigma(\Lambda_c) \cdot 1.14$$

$$\begin{aligned} \text{ALICE: } \sigma_{\text{all}}(d\sigma_{cc}/dy | |y| < 0.5) \\ = C * [\sigma(D^0) + \sigma(D^+) + \sigma(D_s) + \sigma(\Lambda_c) + 2 * \sigma(\Xi_c^0)] \end{aligned}$$

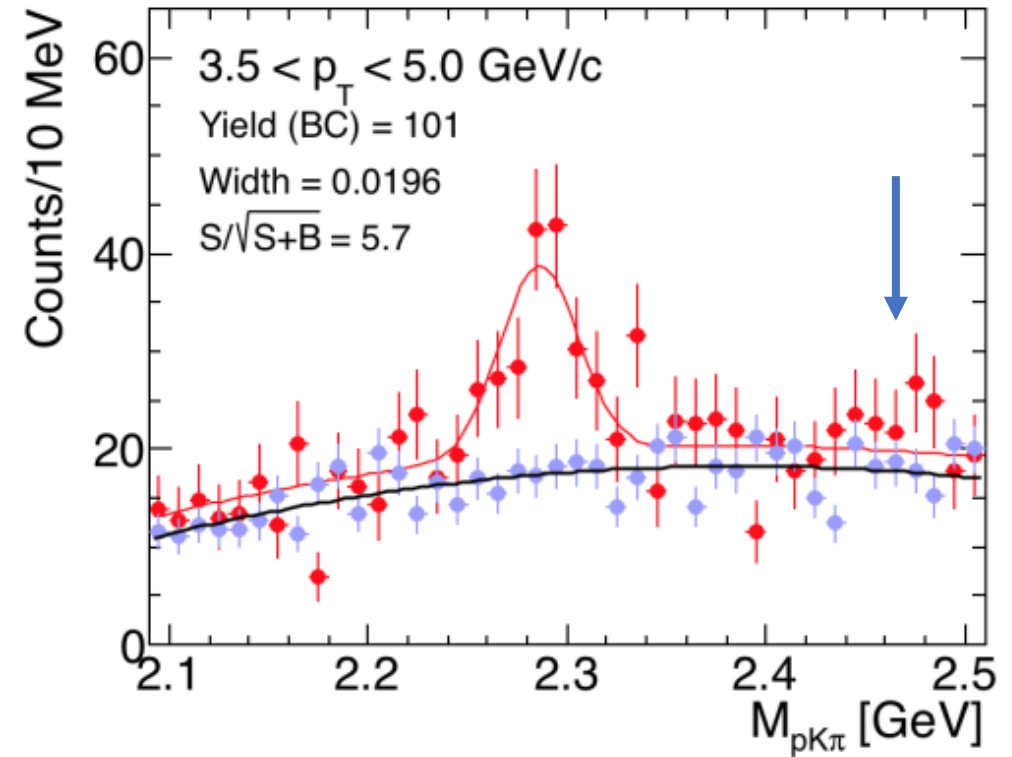
- Ξ_c^0 have a ~10% contribution at ALICE and can not be measured at STAR due to low statistics

possible Ξ_c^+ contribution at STAR

Star run14 AuAu 200 GeV



Star run16 AuAu 200 GeV



Ξ_c^+ Mass $2467.71 \pm 0.23 \text{ MeV}$ $\Xi_c^+ \rightarrow p K \pi$ B.R. $6.2 \pm 3.0 \times 10^{-3}$